

”Antisèche” formulae for simple models

1 Generic properties of Fourier Transform

- **linearity (addition):** $FT[f + g] = FT[f] + FT[g]$,
- **translation (shift):** $FT[f(x - x_0, y - y_0)] = FT[f](u, v) \times e^{2i\pi, (ux_0 + vy_0)}$,
- **similarity (zoom and shrink):** $FT[f(\alpha x, \beta y)] = \frac{1}{\alpha\beta} FT[f](\frac{u}{\alpha}, \frac{v}{\beta})$,
- **convolution (“blurring”):** $FT[f \otimes g] = FT[f] \times FT[g]$,
- ∞ **limit (“small” details):** $FT[f] \xrightarrow{\infty} 0$,
- 0 **limit (“large” details):** $FT[f] \xrightarrow{0} 1$.

2 Generic models

Shape	Brightness distribution	Visibility
Point source	$\delta(\vec{x})$	1
Background	I_0	$\begin{cases} 1 & \text{if } \rho = 0 \\ 0 & \text{otherwise} \end{cases}$
Binary star	$I_0 [\delta(\vec{x}) + R\delta(\vec{x} - \vec{x}_0)]$	$\sqrt{\frac{1+R^2+2R\cos(\frac{\vec{p}\cdot\vec{x}_0}{\lambda})}{1+R^2}}$
Gauss	$I_0 \sqrt{\frac{4\ln(2\varnothing)}{\pi}} \times e^{-4\ln 2 \frac{r^2}{\varnothing^2}}$	$e^{-\frac{(\pi\varnothing\rho)^2}{4\ln 2}}$
Uniform disk	$\begin{cases} \frac{4}{\pi\varnothing^2} & \text{if } r < \frac{\varnothing}{2} \\ 0 & \text{otherwise} \end{cases}$	$\frac{2J_1(\pi\varnothing\rho)}{\pi\varnothing\rho}$
Ring	$\frac{1}{\pi\varnothing} \delta\left(r - \frac{\varnothing}{2}\right)$	$J_0(\pi\varnothing\rho)$
Exponential	$e^{-k_0 r}, k_0 \geq 0$	$\frac{k_0^2}{1+k_0^2\rho^2}$
Any circular object	$I(r)$	$2\pi \int_0^\infty I(r) J_0(2\pi r \rho) r dr$
Pixel (image brick)	$\begin{cases} \frac{1}{lL} & \text{if } x < l \text{ and } y < L \\ 0 & \text{otherwise} \end{cases}$	$\frac{\sin(\pi x l) \sin(\pi y L)}{\pi^2 x y l L}$
Limb-darkened disk (linear)	$\begin{cases} I_0[1 - u_\lambda(1 - \mu)] & \text{if } r < \frac{\varnothing}{2} \\ \mu = \cos(2r/\varnothing) \end{cases}$	$\begin{cases} \left[\frac{\alpha J_1(x) + \beta \sqrt{\pi/2} J_{3/2}(x)}{x^{3/2}} \right]^2 \\ \left(\frac{\alpha + \beta}{2} \right)^2 \\ \alpha = 1 - u_\lambda \\ \beta = u_l \lambda \\ x = \pi \theta_{LD} \frac{B}{\lambda} \end{cases}$

\varnothing = either the diameter for a ring or uniform disk, or FWHM. r or \vec{s} represent the angles in the image plane. $\rho = \frac{B}{\lambda}$ or $\vec{\rho}$ are the spatial frequencies.