



The Navy Precision Optical Interferometer

Current Status, Future Upgrades

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The NPOI Team

■ USNO

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- Mike DiVittorio
- Bob Zavala

■ AZ Embedded Systems

- Tim Buschmann
- David Allen

■ Lowell Observatory

- Jeff Hall
- Gerard van Belle
- Bill DeGroff
- Lisa Foley
- Jason Sanborn
- Susan Strosahl
- Steve Winchester
- Ron Winner

■ NRL

- Richard Bevilacqua
- Sergio Restaino
- **Tom Armstrong**
- Jonathan Andrews
- Ellyn Baines
- Jim Clark
- Henrique Schmitt
- Chris Wilcox

■ TSU

- Matt Muterspaugh
- Mike Williamson

■ Vanderbilt

- Victor Garcia

■ NMT

- Anders Jorgensen



Tour: Navy Precision Optical Interferometer

Flagstaff, AZ (~7,200' / 2,200 m)
79-m baseline
Six 12-cm apertures

[432-m baseline]
[Four 1.8-m apertures]

Anderson
Mesa



Basics



- **NPOI = Navy Precision Optical Interferometer**
 - Major funding by Oceanographer of the Navy and Office of Naval Research
 - Additional instrument funding from National Science Foundation
- NPOI is collaboration b/w USNO, NRL & Lowell Observatory
- Lowell is both a science partner, and a contractor to USNO (infrastructure & ops) & NRL (site projects)



NPOI Current Capabilities

- Simultaneous, group-delay fringe tracking on multiple baselines (6 stations)
- Bandpass 550-850nm in 16 channels ($R \sim 30-50$)
- Single-baseline fringe tracking to $m_V = 6.7$
- Multi-baseline fringe tracking w/closure phase to $m_V \sim 6.0$
- Wide-angle astrometry with 5-10mas accuracy on bright stars
- Operated by one observer, scheduled ~ 355 nights/year

NPOI Imaging Siderostat



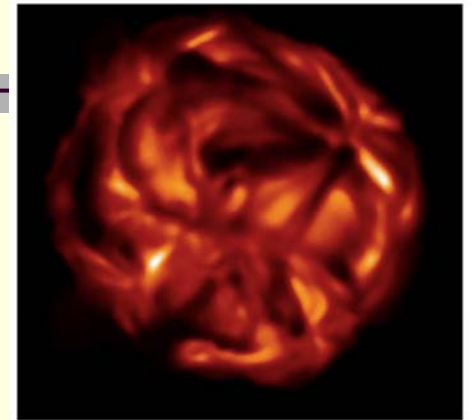
Facility Upgrades in Progress

- Completion of 6-station “imaging” (portable) siderostat array:
 - New enclosures for star acquisition & tip-tilt optics installed for 5 of 6 stations
 - New domes installed for 5 of 6 imaging siderostats
 - 2 more imaging stations to be commissioned in 2013
 - Long baselines to 432 m & ‘compact’ configurations
 - To complete: integration of Long Delay Lines
- Control systems upgrades:
 - PC-based siderostat controllers for astrometric & imaging stations (4 installed; 5 more this year)
 - PC-based Fast Delay Line (FDL) control system:
 - Delivered to site & tested with 2 FDLs thus far
 - Significant performance improvement



'Classic' Instrument Upgrade in Progress

- New Fringe Engine read-out electronics
 - FPGA hardware finished; firmware & software under development
 - 32 → 96 channels, 30^{sec} → unlimited integration time
 - Flexible FTK algorithms
- Baseline bootstrapping, coherent integration
- Recently approved as a 3-year NSF-funded upgrade program



*Freytag simulation
(Chiavassa+ 2010)*



*6x6 pixel imaging
(NPOI 2014?)*



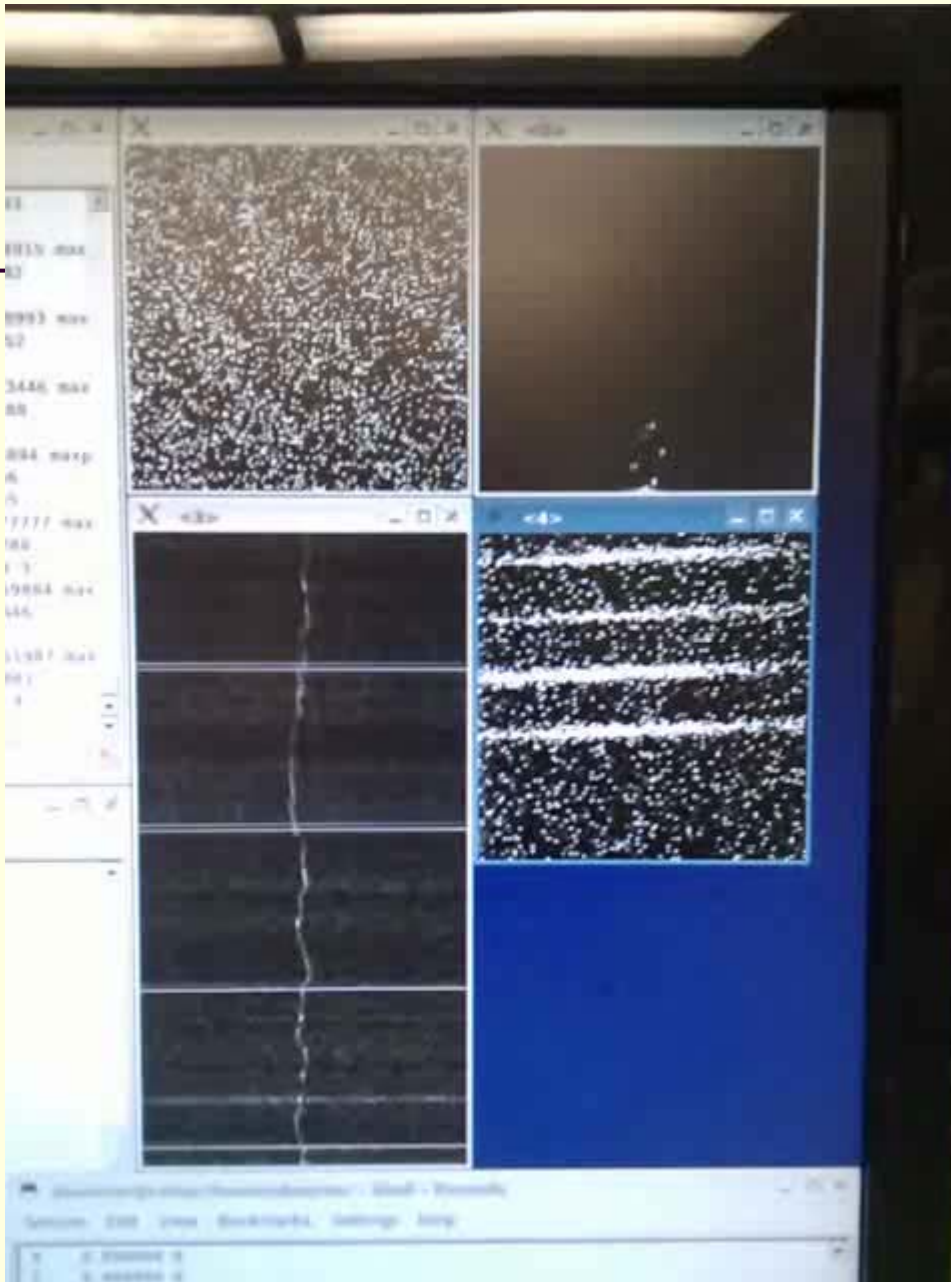
VISION Instrument Commissioning in Progress

- VISION NSF project: PI Matthew Muterspaugh (TSU)
 - Visible “MIRC++” combiner for NPOI
- Improvement
 - Visibility precision $\sim 10\times$
 - No delay nonlinearities, APD afterpulsing
 - Full 6-way combination, flexible spectral resolution ($R\sim 50, 2000$)
 - Modern equipment, CCD
- Commissioning ongoing
 - First fringes: Oct 2012
 - Four-way combination: Jan 2013
 - Grad student (V. Garcia) now on-site for full-time commissioning work



Fringes in Image Plane
(barely seen above the noise)

'Waterfall' plot for single tracking baseline
(left-right motion due to residual atmospheric piston)



FFT of Fringes

- Spatial Frequency versus Delay**
- Each 'Hot Pixel' row is one baseline of six
 - Bottom pixel is DC term
 - Some baselines very hard to see because fringe contrast is low (for long baselines)
 - Left-right motion due to atmospheric piston
 - Pixels should line up in a single vertical line but don't due to uncorrected static piston offsets

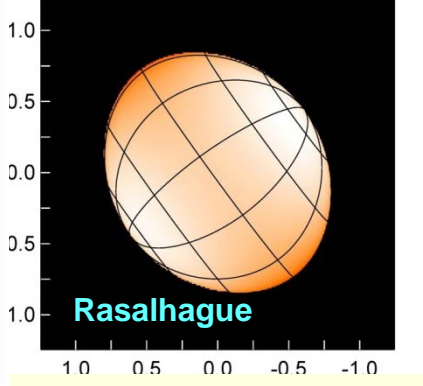
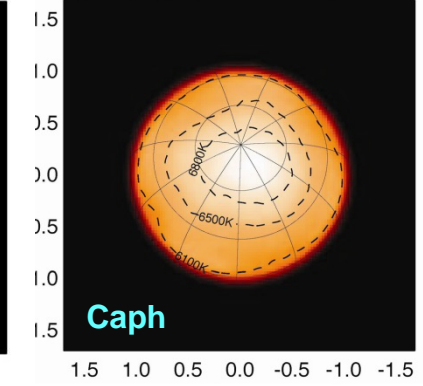
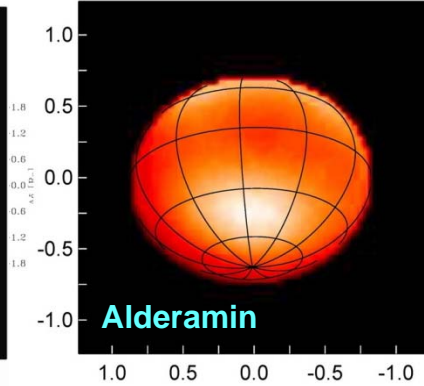
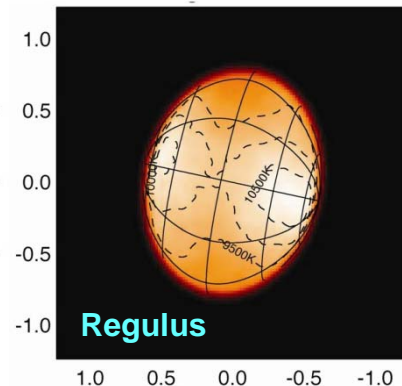
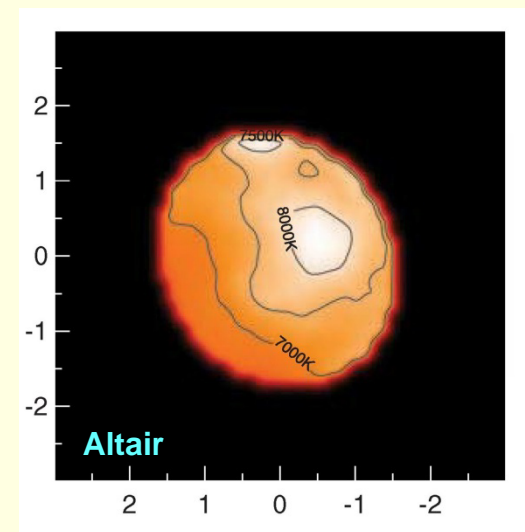
Photometry for each of 4 telescopes

**Four simultaneous telescopes
=
Six baselines
(and 3 independent closure phases)**



NPOI Targets: Stars are Photogenic

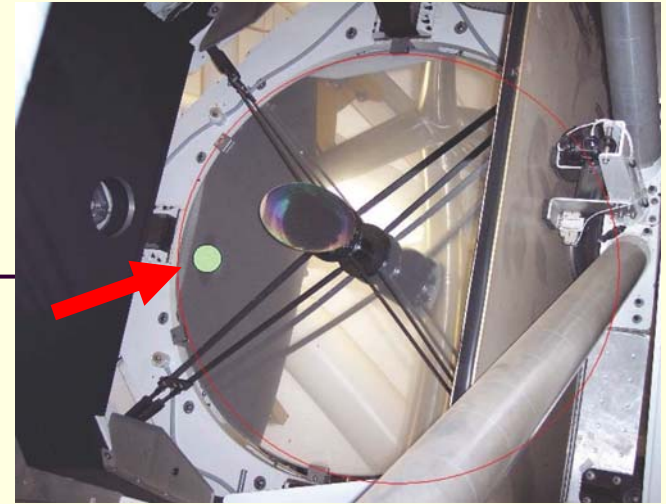
- Progression over the past 10 years
 1. Simple modeling
 2. Detection of surface features
 3. Direct imaging
- Already starting to see some surprises
 - Stellar structure not as expected from simple models
- Next step: time-series images → movies



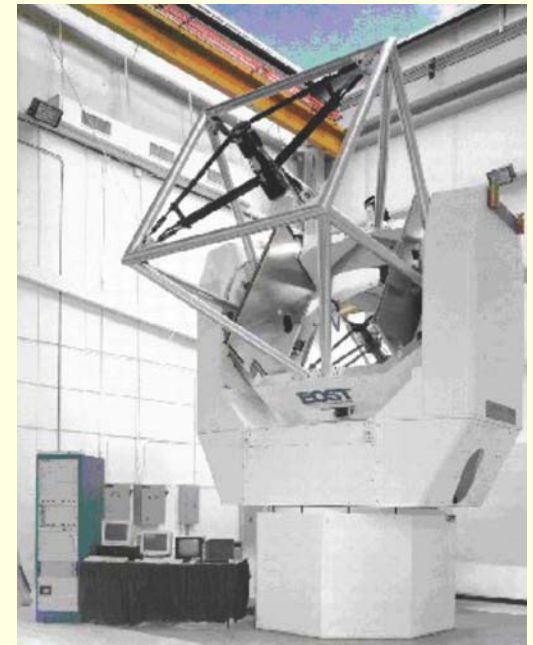
Monnier+ 2007, Zhao+ 2009, Che+ 2011: 4-way CHARA-MIRC Images

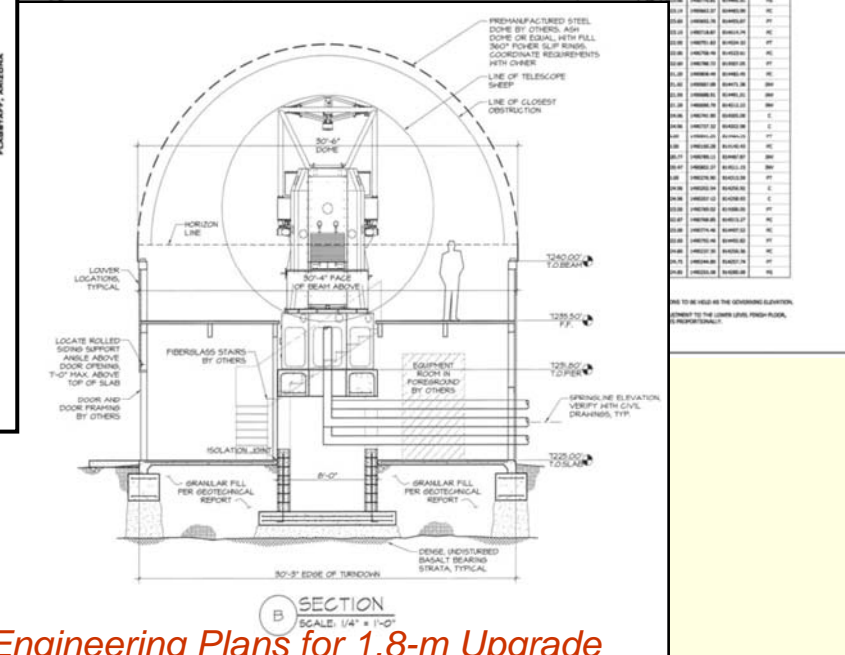
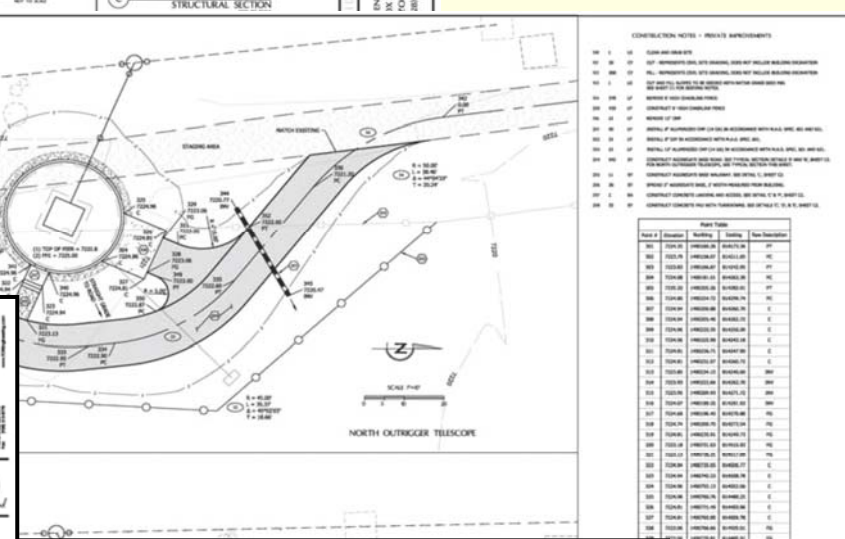
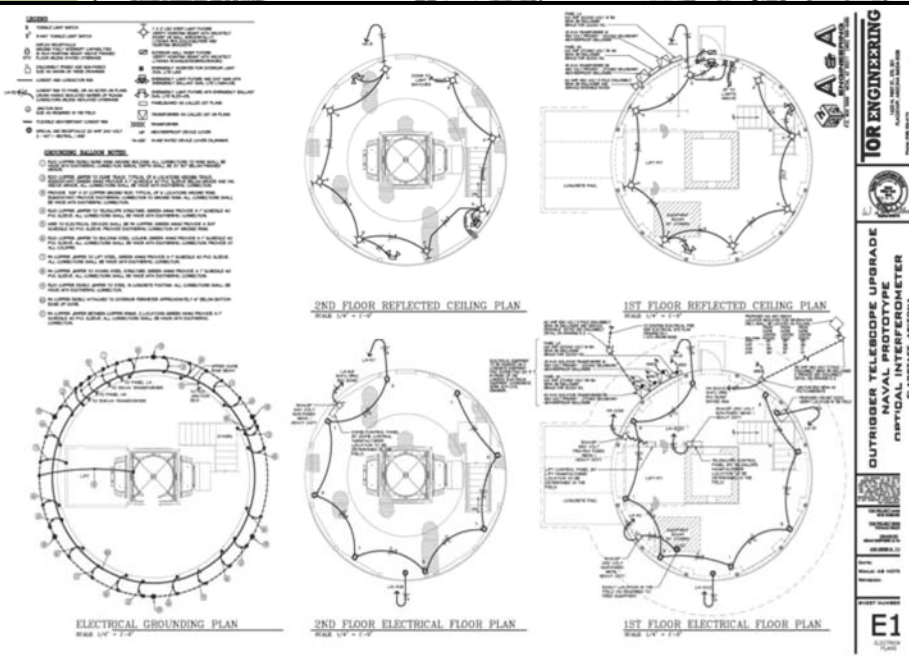
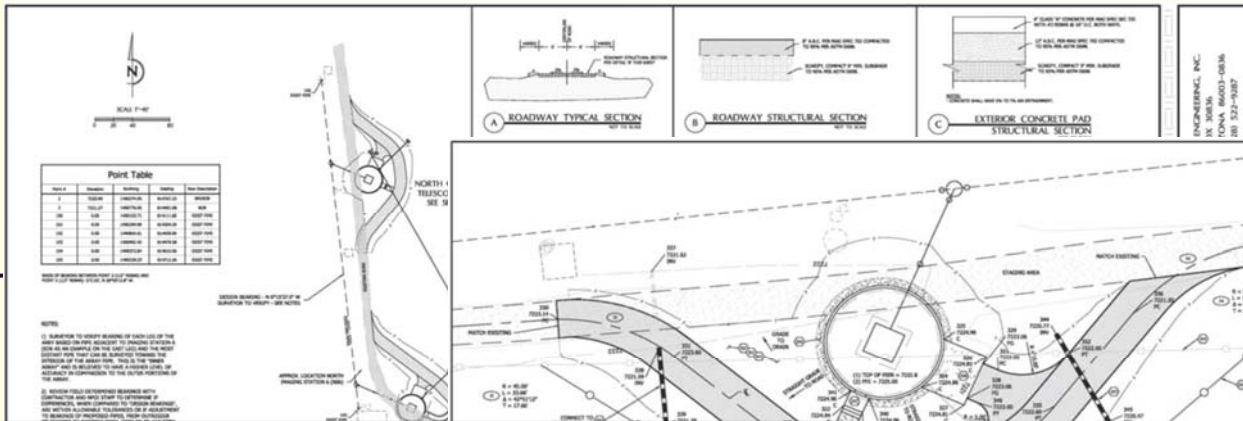
1.8-m Upgrade

- Four 1.8-m telescope were built by NASA for the Keck Interferometer
 - Not installed for non-technical reasons
 - 'Over-engineered' for narrow-angle astrometry
 - Good for NPOI wide-angle astrometry mission
 - [One on loan to Mt. Stromlo]
- 'Gifted' to USNO from CARA in 2010
- Final engineering plan in 2011
- USFS site permit in 2012



Relative size of current 5" siderostats and 1.8-m telescope





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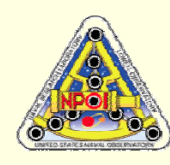
NORTH AND CENTER OUTRIGGER TELESCOPES
GRADING AND DRAINAGE
NAVAL PROTOTYPE OPTICAL INTERFEROMETER
FLAGSTAFF, ARIZONA

PRELIMINARY
NOT FOR CONSTRUCTION

J. W. LOWELL
100-700-0000
100-700-0000
100-700-0000

DATE: 10/10/2011
TIME: 10:00:00
USER: J. W. LOWELL
PROJECT: NAVAL PROTOTYPE OPTICAL INTERFEROMETER
FILE: 100-700-0000

Prepared Engineering Plans for 1.8-m Upgrade



Ready to be
done with the
VLTI school?

TODAY'S LESSON :

1. Ongoing upgrades
 1. VISION
 2. Classic
 3. Additional stations
 4. Long baselines
2. 1.8-m apertures are coming!

Any questions?

