


# **EXors – cism with AMBER and MIDI**

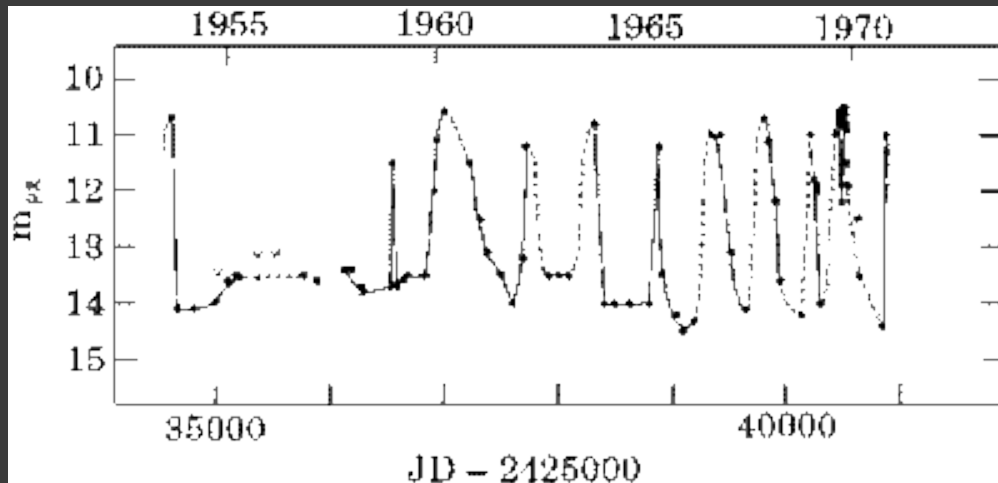
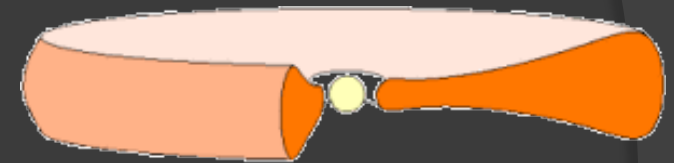
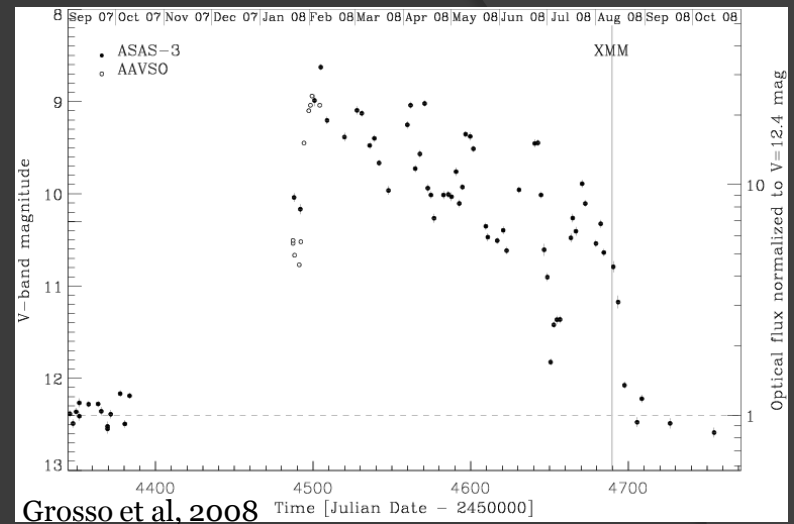
(blindly measuring stars when they are the most boring (probably))



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# Young Stellar Objects

- Eruptive T Tauri stars
- Prototype: EX Lupi 
- 1 – 4 mag outbursts over 10 – 100 days
- Outbursts separated by several months
- Eruptions attributed to enhanced accretion rate
- Bry line is good indicator of accretion rate (Lorenzetti et al., 2012)



Light curve of the EXor-type star VY Tau. (From G.H. Herbig: *Eruptive phenomena in early stellar evolution*, *Astrophysical Journal* 217, 1977, pp. 712)



V1515 Cyg

# Targets

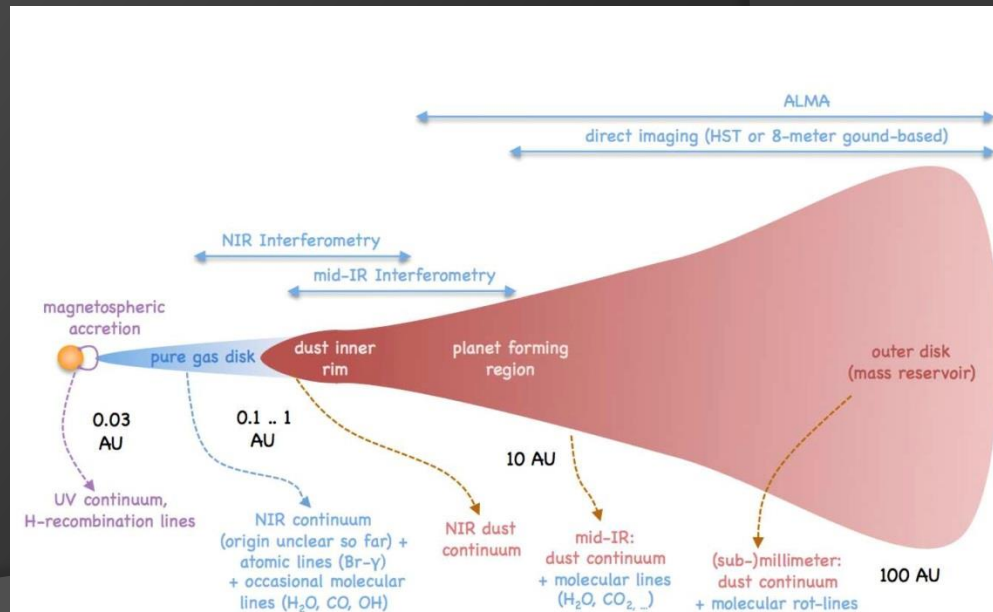
DR Tau – K=6.9, N=3 140 pc  
- archival MIDI, Br $\gamma$  detection



NY Ori – K=8, N=3.3 400 pc



V2775 Ori – K=11.8, N=3  
400 pc



Dullemond & Monnier, 2010

# Observations

Zero Epoch observations

DR Tau – AMBER and MIDI  
(1+1)

NY Ori – AMBER (?) and MIDI  
(1+2)

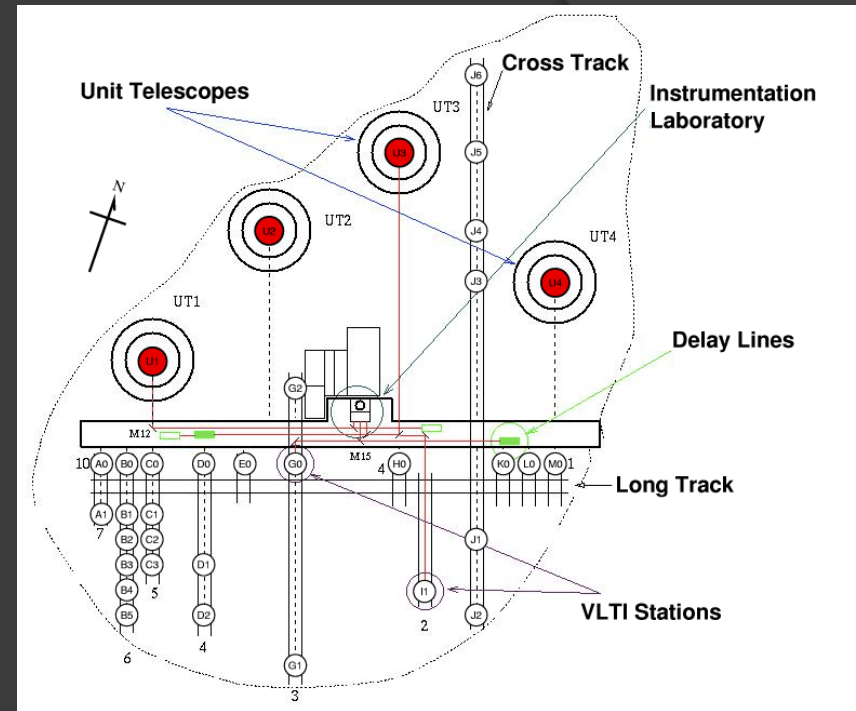
V2775 Ori – MIDI  
(2)

MIDI – UT2-UT3 46 m baseline (DR Tau)

UT2-UT3 and UT2-UT4 89m baseline (NY Ori, V2775 Ori)

AMBER – UT1-UT3-UT4 (DR Tau)

UT1-UT2-UT3 (NY Ori) – fainter star

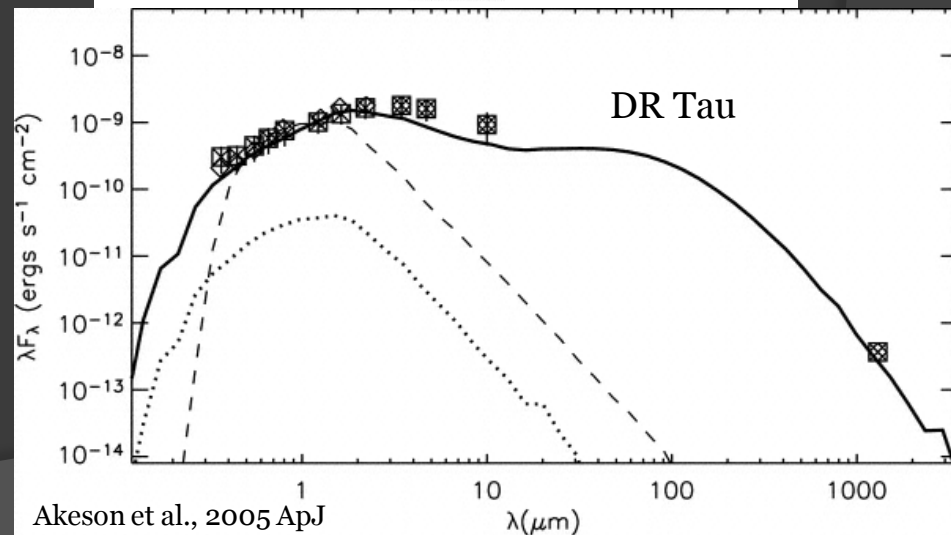
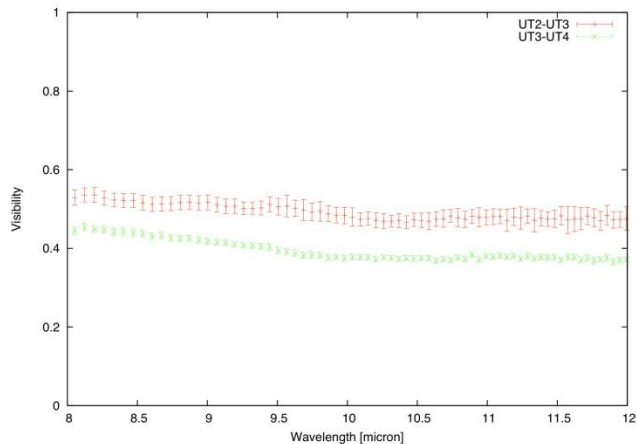
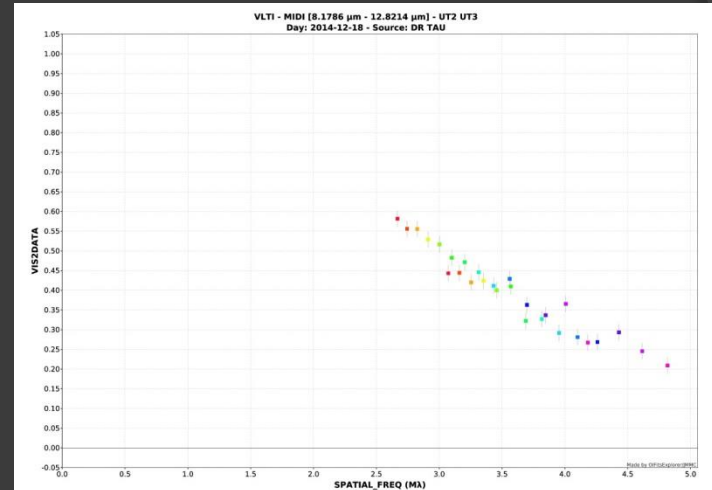
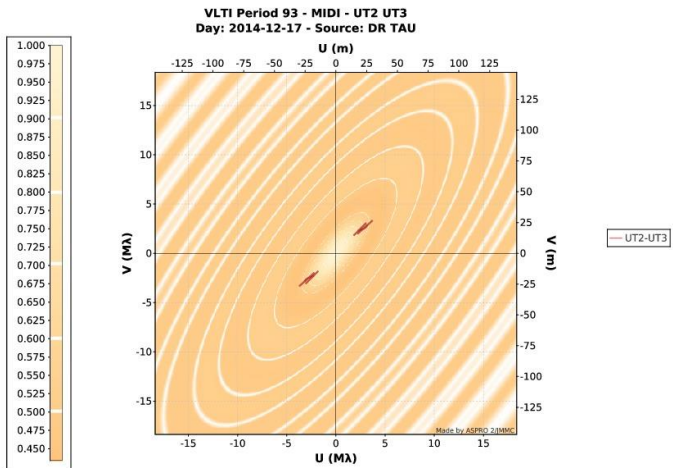


Ask for ToO when outbursting

# EX-pectations

Bary line measurement from AMBER will give accretion rate for DR Tau

Inner and outer disk size measurements for all stars with AMBER and MIDI



# References

[http://kisag.konkoly.hu/Science/eruptive\\_intro.html](http://kisag.konkoly.hu/Science/eruptive_intro.html)

<http://inspirehep.net/record/861675?ln=en>

**Akeson et al., 2005 ApJ**

**Dullemond & Monnier, 2010**

**Grosso et al., 2008**

**Herbig, G., 1977 ApJ**

