DIRECT OBSERVATION OF EXOPLANETS WITH GRAVITY

VLTI SCHOOL 2021 S. LACOUR

ON THE DIFFICULTY OF DIRECT DETECTION











Bern model of planet formation and evolution (Mordasini et al. 2012) Including the Young Stellar formation regions (Cha, Sco-Cen) Including Young moving groups (Beta Pictoris, TW Hya, Carina, etc..)



Stellar light (Extreme AO, 95%Strehl)







500 mas

EC=58/60%;t0=8.1/9.7ms;texp=420/420s





Fomalhaut b Planet 2006 17/06/2021 9



17/06/2021

THE GRAVITY INTERFEROMETER





GRAVITY @ ESO VLTI





The GRAVITY cryostat



Haug, Thiel, Hausmann & the GRAVITY collaboration











Haug, Thiel, Hausmann & the GRAVITY collaboration

FIBERCOUPLER



FIBERCOUPLER







 $\delta OPD = \vec{B} \cdot (\vec{\alpha} - \vec{\beta})$



	Metrology phases	Fringe Tracker visibilities
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Science camera visibilities

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	PRIMARY		IMAGE	1 x 1	load						
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	IMAGING_DATA_SC		IMAGE	360 x 288 x 30	load)					
	IMAGING_DETECTOR_SC		BINART TABLE	10 X 24	REGION	DETECTOR	PORTS	CORRELATION			
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Cancel

Read All



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RESULTS FROM OPTICAL INTERFEROMETRY









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Nowak et al. (2020) Lagrange et al. (2020) → Press release tomorrow







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4T





5
2

Parameter	Prior distribution	Poster	Unit		
Star		βF			
Stellar mass	Gaussian (1.77 ± 0.02)	1.82	M_{\odot}		
Parallax	Uniform	51.42	\max		
proper motion	Uniform	$\texttt{ra}{=}4.88\pm0.02$	m mas/yr		
$RV v_0$	Uniform	-23.0	m/s		
RV jitter	Uniform	31.2	1 ± 8.09	m/s	
Planets		β Pictoris b	β Pictoris c		
Semi-major axis	Log Uniform	9.90 ± 0.05	2.72 ± 0.02	au	
Eccentricity	Uniform	0.10 ± 0.01	0.37 ± 0.12		
Inclination	Sin Uniform	88.99 ± 0.01	89.17 ± 0.50	\deg	
PA of ascending node	Uniform	31.82 ± 0.02	30.98 ± 0.12	\deg	
Argument of periastron	ent of periastron Uniform		196.9 ± 3.5 66.2 ± 2.5		
Epoch of periastron	ch of periastron Uniform		0.83 ± 0.02		
Planet mass	Gaussian (15.4 ± 3.0)		9.0 ± 1.6		
	Log Uniform		8.2 ± 0.8	M_{jup}	
Period		23.28 ± 0.46	3.37 ± 0.04	years	
$\Delta m_{\rm K}$		8.9 ± 0.1	10.8 ± 0.1	mag	

Nowak et al. (2020)

PDS 70 B & C

PDS 70 B & C

Gaussian prior on coplanarity with disk : Sigma = 10 degrees

Gaussian prior on coplanarity between b & c

Wang et al. (work in progress)





2000 Rjup

 \rightarrow



With circumplanetary model, for PDS70b: Rcpd < 1mas (0.1au)

> R Hill = 1.6 au R Bondi = 8 au



K BAND SPECTRA (J. WANG ET AL. IN PREPARATION)





- PLANETARY-TYPE ATMOSPHERE
- NO BLACK BODY CPD
- NEED TO ADD EXTINCTION

BETA PICTORIS B

- 23 MYR OLD
- 20 PC
- ~ 12 MJUP PLANET



Beta Pictoris b



Gravity Collaboration: M. Nowak et al. (2020)

Spectral Modeling

Self-consistent Chemical Modeling Cloud Modeling Radiative-convective temperature => "correct" / slow / can use sparse data

Free retrieval

Abundance Cloud opacity Temperature Parametrize! => "data speak" / fast / compare models

- complicated
- unknown
- love to blow up self-consistent codes





Planet composition varies with disk radius



This simple picture ignores chemistry (freeze out only)





PROSPECTS FOR OPTICAL INTERFEROMETRY





8





The Very Large Telescope in 2030 – White paper proposed to ESO – strongly supported

GRAVITY⁺ : Towards foint science, all sky milliarcsec optical interferometric imaging

STC recommends immediate start

GRAVITY⁺: Towards faint science, all sky milliarcsecond optical interferometric imaging

Constructions by Performing and the Construction of the Constructi

Introduction

Ever since the development of 'aperture synthesis' spatial interferometry in the raband in the 1950s (Syr& Hewshi 1946), this technique has been the standard choice for development of large aperture, high angular resolution tolescopes in the microwave a radio bands (Thompson, Moran & Swenson 1986).

The probability of radia interference type has based on assuming the range has much shifting the same based on the same set of the same set of the same set of the same set of the same based on the same set of the same based on the same set of the same s

> s over a wide field-of-view. evel polarimetric m extral astrometry at the level of 0.06 degrees rms in interfer



Off Axis Tracking

Adaptive Optics

Laser Guide Stars

Improved Sensitivity

ExoGRAVITY : exoplanets are design-driver science case

Credit: ESO, Huedepohl