

The background of the slide is a dark, star-filled space. In the upper left, a bright yellow star is visible. In the center-right, a large, reddish-brown gas giant planet with horizontal bands is shown. In the lower-left foreground, a smaller, dark brown dwarf is partially visible, overlapping the larger planet.

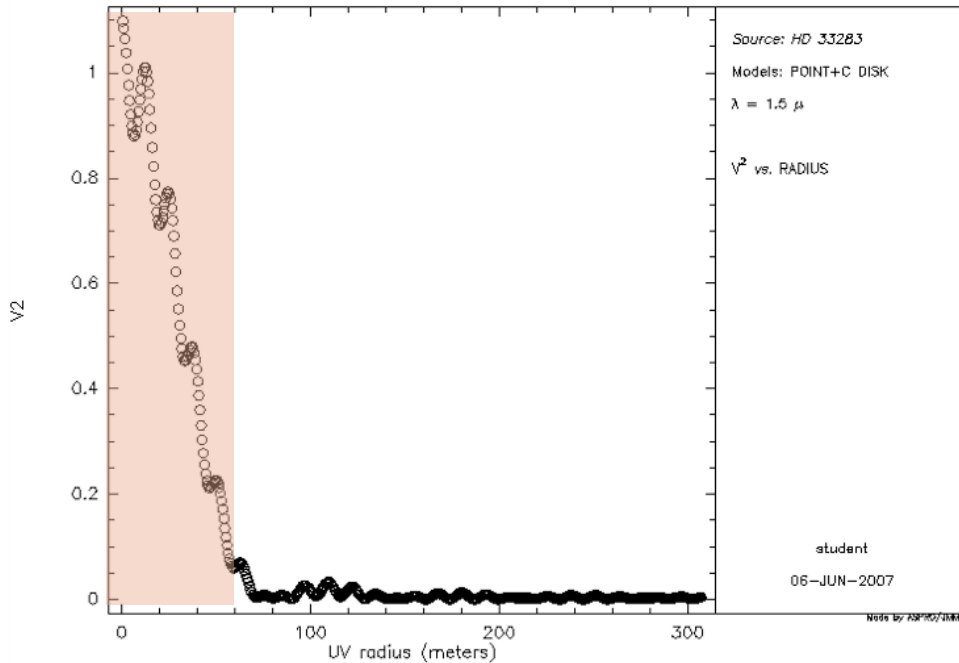
The Search for Brown Dwarfs among Exoplanet Candidates

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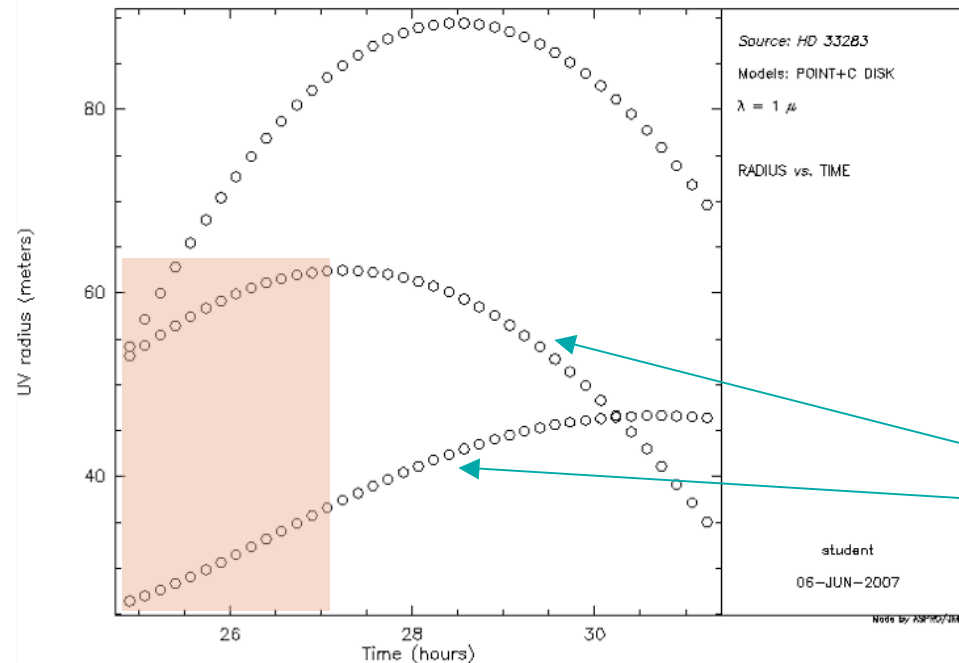
Scientific Motivation

- Learn more about exoplanets and their environments
 - Radial velocities give **a** and **$M \sin i$**
 - 209 Candidates found
 - 6% chance 1 object (with $M \sin i < 1 M_J$) has $M > 13 M_J$
 - Use VLTI to determine the photometric properties
- Search for E.T.

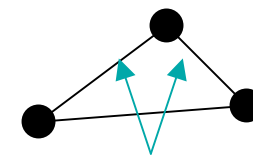


- Observational requirements:
- Short baselines
 - Close to source rise/set
 $\Delta_{\text{bsl}}/\Delta_{\text{time}} \sim 1.2\text{m/exptime}$
 for typical separation (ideal)
 - Low contrast
 (errors on visibility set contrast limit of $<1\%$ for one observation)

Plot assumptions:
 stellar diameter $\sim 5\text{mas}$
 median separation $\sim 24\text{mas}$
 dec = -26°



Instrument: AMBER
 With UT 2, 3, 4



Perpendicular baselines

Minimum Detectable Flux Ratio Drives Target List

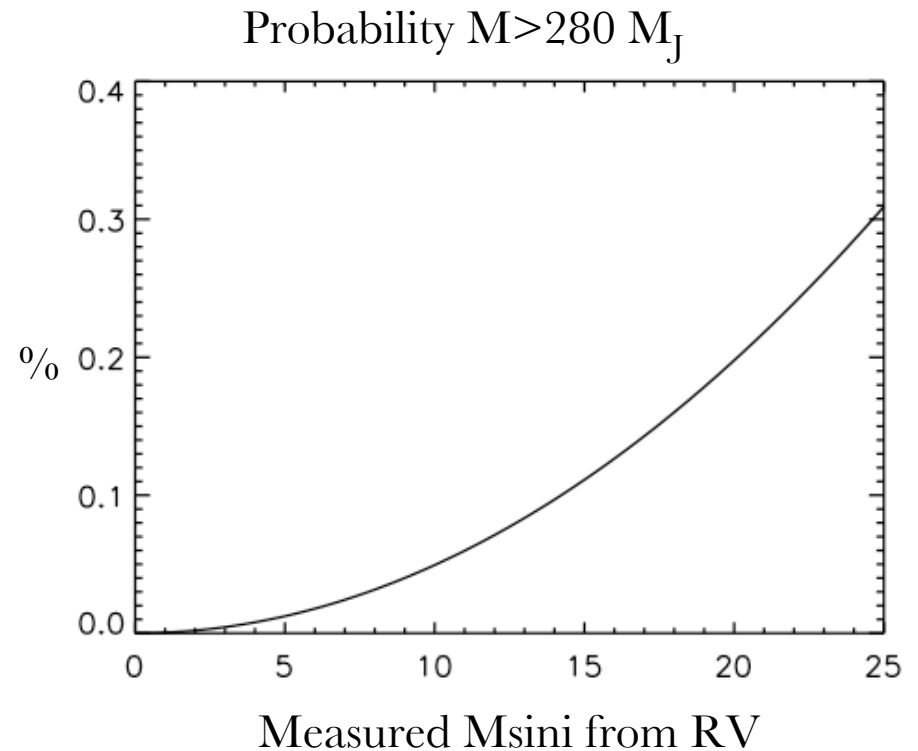
- median J-mag of Host Stars ~ 6.5
- VLTI gives $\sim 1\%$ V^2 errors
- 1/100 min flux ratio

$$\frac{1}{100} = \frac{M_{bd}^{3.5}}{M_*^{3.5}}$$

- Assume Solar Mass Host Star
- Minimum Detectable Target Mass:

280 M_J !!!

(M Dwarf)



Only $M_{sini} > 5 M_J$ have chance of being resolved

~ 30 such objects

3% chance one $280 M_J$ will be detected!

Compare χ^2 for V^2 and
what do we expect to find?



M dwarfs



Brown dwarfs



(White dwarfs?)