#### **Tutorial T-10 Amber Data Reduction (practical point of view)**

**EuroSummer School** 

**Observation and data reduction with the Very Large Telescope Interferometer** 

Goutelas, France June 4-16, 2006

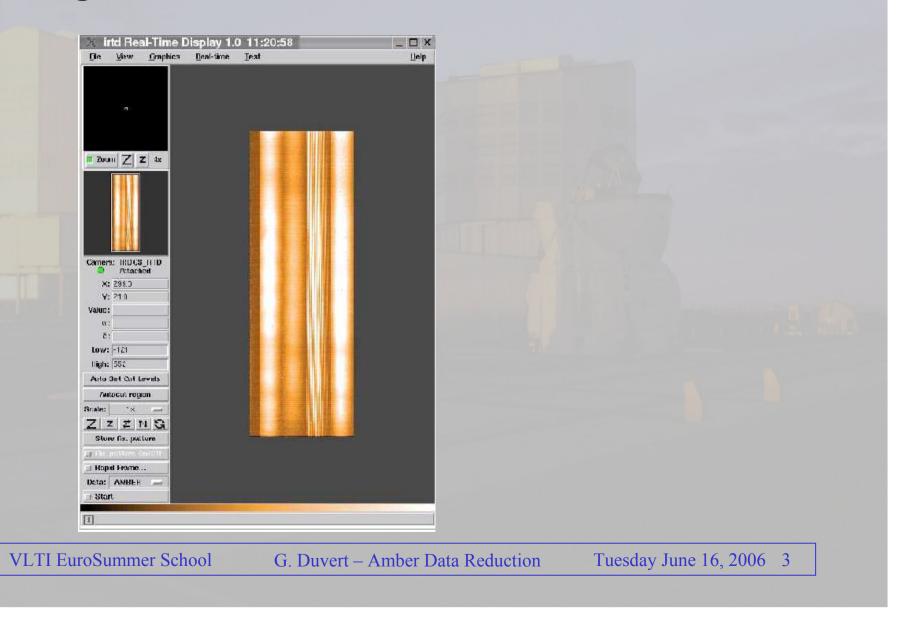
G. Duvert LAOG/JMMC 13 June 2006

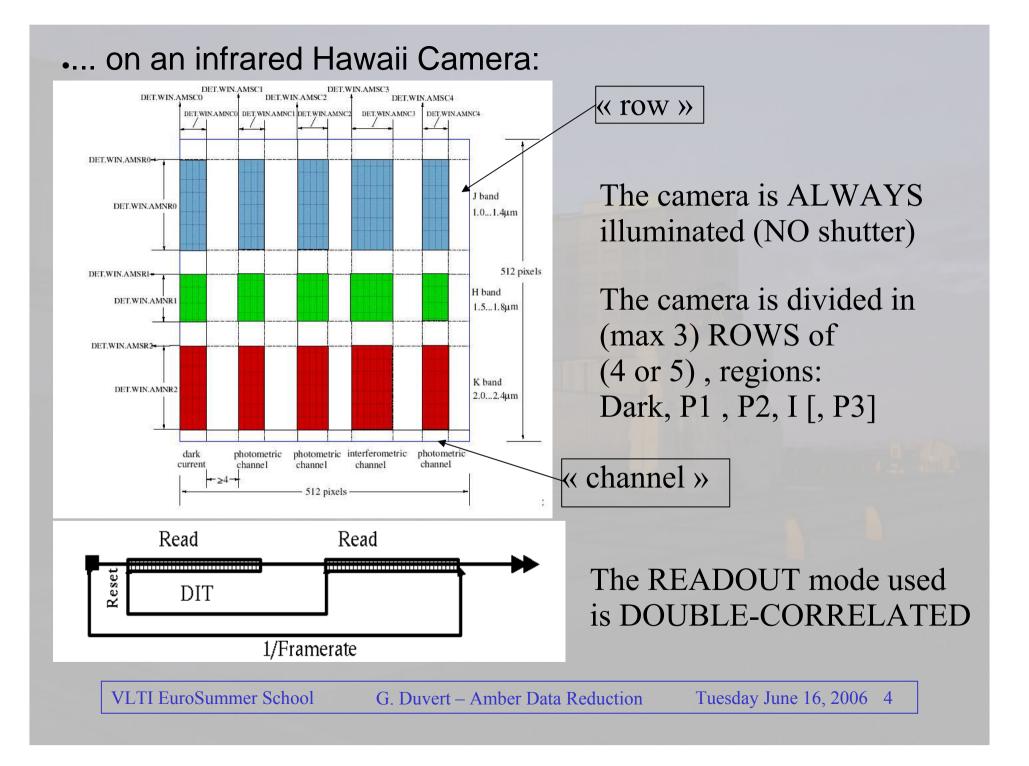
## AMBER DATA REDUCTION OVERVIEW

AMBER paradigm: *spatially coded, spectrally dispersed, photometrically monitored, fringes.* 

- . fringes ... on an infrared Hawaii Camera:
- camera readout mode
- camera windowing, readout timing
- . camera readout noise, bad pixels, flat, etc...
- spectrally dispersed ... needs spectral calibration
- wavelength calibration
- spatially coded ... needs spatial coding calibration:
- the P2VM

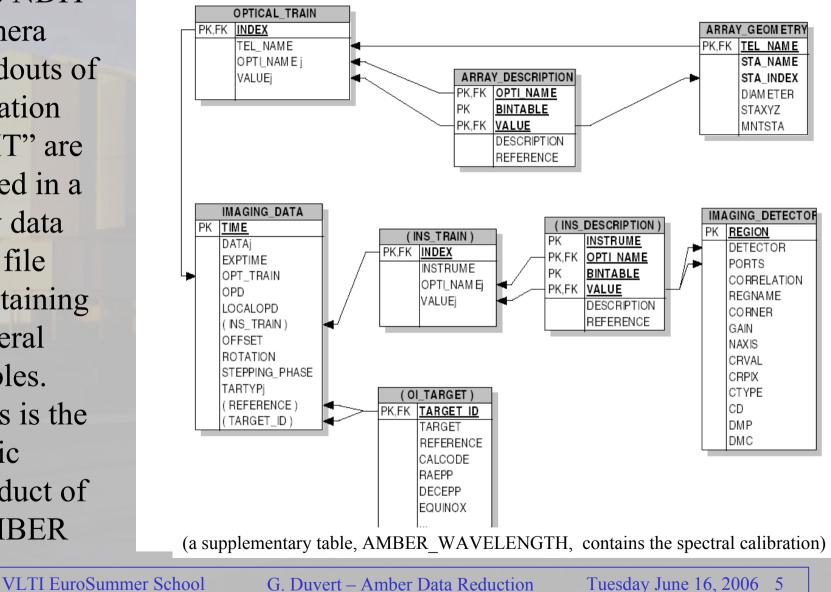
#### . fringes ...



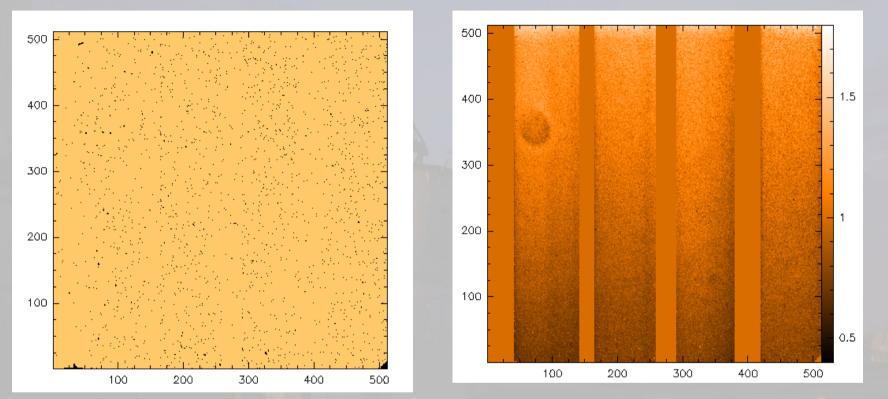


#### RAW DATA Fits file: sequence of FITS Tables, described in ESO Doc No · VLT-SPE-ESO-15000-2764

The NDIT camera readouts of duration "DIT" are saved in a raw data fits file containing several Tables. This is the basic product of AMBER

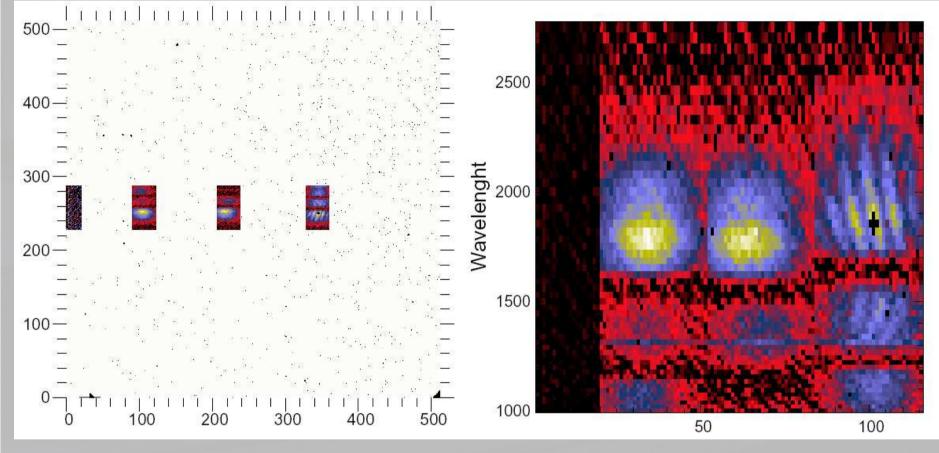


#### . camera readout noise, bad pixels, flat, etc...



- 1. Bad Pixels -> "Bad Pixel Map" File
- BIAS depends on the illumination of the camera and EXPOSURE TIME -> "Dark" Files
- 3. Relative pixel-to-pixel gain -> "Flat Field Map" File

#### spectrally dispersed ...



...displacement of photometric « channels » - CALIBRATE – 3 or 4 Files ...accurate wavelength calibration of the Interferometric « channel »? NO

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Tuesday June 16, 2006 7

## • spatially coded ... the P2VM: 5 (2T) or 9 (3T) files

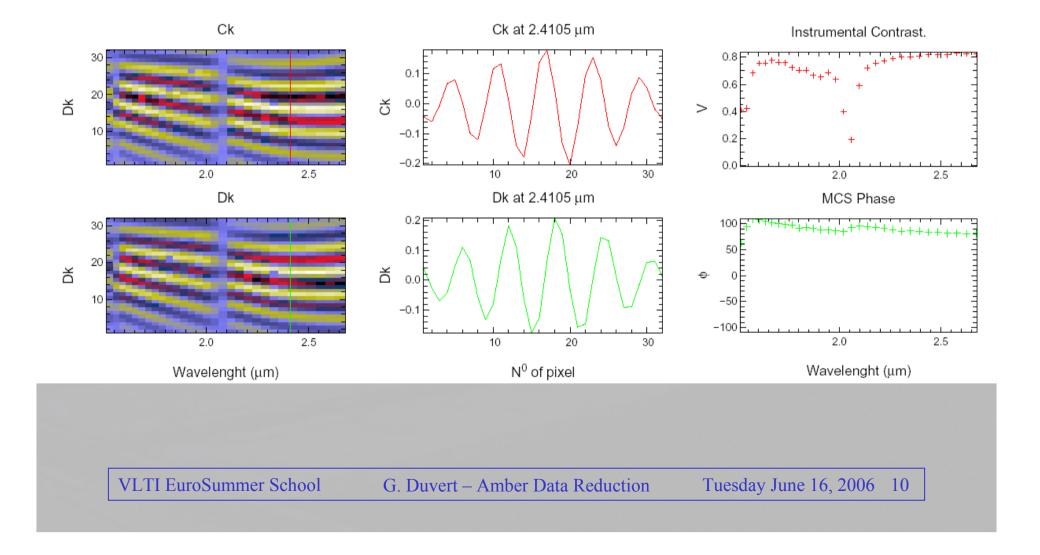
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Open	Open	Close	1/2 Delayed	AMBER_3TSTD_CAL_0005.fits	Notice Photos Tanter of Pool	
		Figure	3. Complete calib	ration sequence for 2 telescopes	·	
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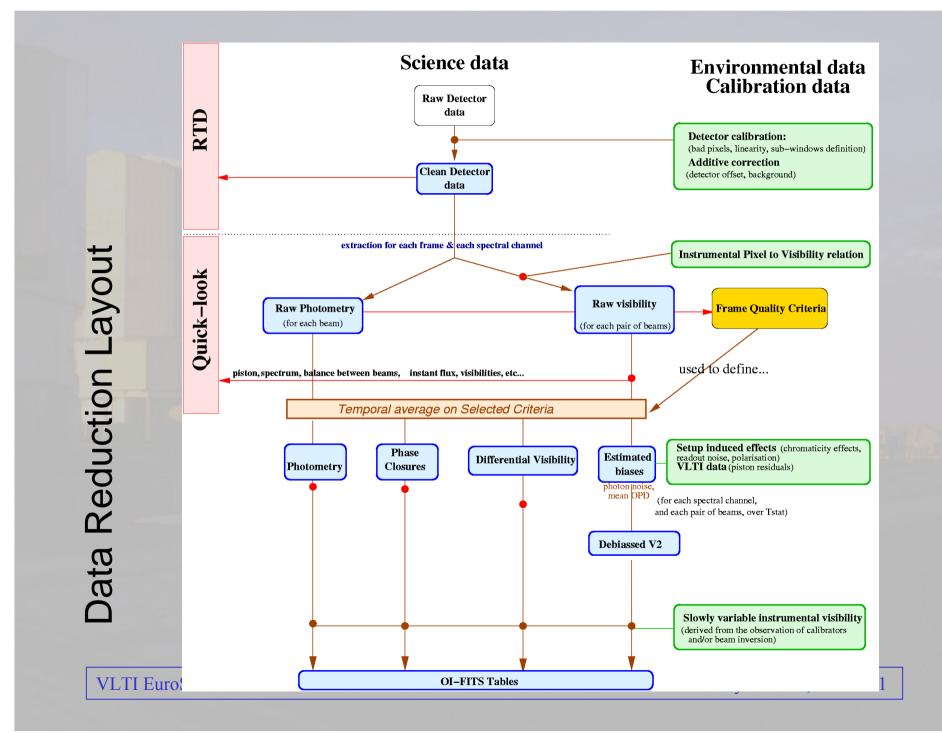
# Shutters and P2VM calibration files

Step	Shutter 1	Shutter 2	Shutter 3	Phase $\gamma_0$	DPR key
1	Open	Closed	Closed	NO	2P2V, 3P2V
2	Closed	Open	Closed	NO	2P2V, 3P2V
3	Open	Open	Closed	NO	2P2V, 3P2V
4	Open	Open	Closed	YES	2P2V, 3P2V
5	Closed	Closed	Open	NO	3P2V
6	Open	Closed	Open	NO	3P2V
7	Open	Closed	Open	YES	3P2V
8	Closed	Open	Open	NO	3P2V
9	Closed	Open	Open	YES	3P2V

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# the P2VM: the « carrying waves: Ck and Dk»





#### Sequence of observations:

File Selected files Tools Help

B 🖑 🙇	Default grouping	expand	Find entry:	▼ find			
File	TPL.ID	DATE-OBS	ORIGFILE	OCS.OBS.SP	OBS.TARG.N]	DPR.CATG	DPR.TYP
0 200147568 HD124454-Hummel-Cal							
AMBER.2005-02-24T04:32:19.375.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_ADJUST_COLPOS055_0003.	Low_JHK	hd124454	CALIB	STD
- AMBER.2005-02-24T04:33:20.472.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0017.fits	Low_JHK	hd124454	CALIB	WAVE, 2TE
- AMBER.2005-02-24T04:33:29.317.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0018.fits	Low_JHK	hd124454	CALIB	WAVE, 2TI
- AMBER.2005-02-24T04:33:41.750.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0019.fits	Low_JHK	hd124454	CALIB	WAVE, 2T
- AMBER.2005-02-24T04:34:28.205.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0020.fits	Low_JHK	hd124454	CALIB	2P2V
- AMBER.2005-02-24T04:34:47.999.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0021.fits	Low_JHK	hd124454	CALIB	2P2V
- AMBER.2005-02-24T04:34:48.999.fits		2005-02-24T	AMBER_P2VM_2T_2005-02-24T0	. Low_JHK	hd124454	CALIB	P2VM
- AMBER.2005-02-24T04:35:00.527.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0022.fits	Low_JHK	hd124454	CALIB	2P2V
- AMBER.2005-02-24T04:35:13.453.fits	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0023.fits	Low_JHK	hd124454	CALIB	2P2V
AMBER.2005-02-24T04:35:26.200.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_ACQ055_0024.fits	Low_JHK	hd124454	CALIB	2P2V
- AMBER.2005-02-24T05:18:33.557.fits.gz	AMBER_2T_cal	2005-02-24T	AMBER_BASE12_CAL055_0023.fits	Low_JHK	hd124454	CALIB	WAVE, 2T
- AMBER.2005-02-24T05:26:48.689.fits.gz	AMBER_2T_cal	2005-02-24T	AMBER_BASE12_CAL055_0024.fits	Low_JHK	hd124454	CALIB	WAVE, 2T
- AMBER.2005-02-24T05:33:02.551.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_DARK055_0007.fits	Low_JHK	hd124454	SCIENCE	DARK
- 🖸 AMBER.2005-02-24T05:33:51.335.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0031.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T05:34:36.498.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0032.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T05:35:22.560.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0033.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T05:36:25.715.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_SKY055_0007.fits	Low_JHK	hd124454	CALIB	SKY
- AMBER.2005-02-24T05:38:13.629.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_DARK055_0008.fits	Low_JHK	hd124454	SCIENCE	DARK
- AMBER.2005-02-24T05:39:32.361.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_085055_0034.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T05:40:52.249.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0035.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T05:42:07.442.fits	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_OBS055_0036.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- 🖬 AMBER.2005-02-24T05:43:38.720.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_SKY055_0008.fits	Low_JHK	hd124454	CALIB	SKY
AMBER.2005-02-24T05:55:59.195.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0037.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T05:57:14.516.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0038.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T05:58:46.843.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_SKY055_0009.fits	Low_JHK	hd124454	CALIB	SKY
- AMBER.2005-02-24T06:01:28.562.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_DARK055_0009.fits	Low_JHK	hd124454	SCIENCE	DARK
- AMBER.2005-02-24T06:02:47.447.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0039.fits	Low_JHK	hd124454	SCIENCE	OBJECT
- AMBER.2005-02-24T06:04:02.198.fits.gz	AMBER_2Tstd	2005-02-24T	AMBER_2TSTD_0BS055_0040.fits	Low_JHK	hd124454	SCIENCE	OBJECT

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Tuesday June 16, 2006 12

A dedicated library, **amdlib**, is pivotal for both observation & data reduction.

It is used at the time of observation to find fringes, etc...

All the steps for data reduction are available as functions in the amdlib library.

These functions can be called from external programs, but amdlib provides also command-line programs to perform the data reduction.

Each raw data file (including those used to obtain calibration files such as the P2VM) must be calibrated for camera cosmetics.

Internally each calibrated file is converted to a simpler structure, called ScienceData:

-The temporal sequence of 2D Interferograms, untouched.

- Each 2D photometric channel is replaced by a 1D photometry vector (1 instantaneous photometry per beam and per wavelength), and associated errors.

#### A P2VM must be obtained before any Science Data; The P2VM File is computed by the amdlib command amdlibComputeP2vm

# Using the 5- or 9- calibration files obtained on the internal source of AMBER

Usage: amdlibComputeP2vm [-s o1,o2,o3] badPixelMap flatFieldMap dark file1 ... file9 p2vmFile

or

amdlibComputeP2vm [-s o1,o2,o3] badPixelMap flatFieldMap dark file1
... file4 p2vmFile

where ol,o2... is a comma-separated list of shifts for the 2 or 3 photometric channels

The DATA necessary to get visibilities on a Science Object are:

- RAW DATA on a Science Object (AMBER\_XXX\_OBS\_YYY)
- Associated Pixel Bias Map (aka DARK)
- [eventually an associated SKY]
- The P2VM in use at time of the observation.

The command amdlibExtractVis does the complete sequence of calibration and visibility, etc., computation using the P2VM, and produces an OI\_FITS file:

```
Usage: amdlibExtractVis [-f] [-s] [-b num] [-p PHASOR|PHASE][-e
STATISTIC|THEORIC] badPixelMap flatFieldMap p2vm dark sky input
output
```

```
Options: -f to force the use of another P2VM
```

- -s split OI-FITS; one file per band
- -b 'num' to average (bin) over 'num' frames
- -p PHASOR or PHASE (default) for piston algorithm
- -e THEORIC or STATISTIC (default) for error bars estimation
- -c FRG or FLUX criterion for frame selection
- -r ratio for frame selection

# A number of programs already call amdlib and provide and interface to it:

-ammYorick, a Yorick implementation by F. Millour (et al).

- a scripting interface in the ESO's gasgano file sorter (to be used this evening)

The ESO pipeline, now to be delivered with a flexible workflow bench, "taverna" and the gasgano interface (next september).
The Final OI\_FITS product can be read and used by many generalist data visualisation programs (IDL, gildas, etc...)

# OI\_FITS Product:

Normalized by the IAU commission and published. (Pauls, T. A.; Young, J. S.; Cotton, W. D.; Monnier, J. D., "A Data Exchange Standard for Optical (Visible/IR) Interferometry", 2005, The Publications of the Astronomical Society of the Pacific, Volume 117, Issue 837, pp. 1255-1262.). Has a number of Extensions:

Index	Extension	Туре	Dimension		Vie	w		
_ 0	Primary	lmage	0	Header	lm	age	1	Fable
⊒1	OI_ARRAY	Binary	5 cols X 3 rows	Header	Hist	Plot	All	Sele
_] 2	OI_TARGET	Binary	17 cols X 1 rows	Header	Hist	Plot	All	Sele
_] <b>3</b>	OI_WAVELENGTH	Binary	2 cols X 158 rows	Header	Hist	Plot	All	Sele
_ 4	OI_VIS	Binary	14 cols X 1500 rows	Header	Hist	Plot	All	Sele
_] 5	OI_VIS2	Binary	10 cols X 1500 rows	Header	Hist	Plot	All	Sele
_] 6	0І_Т3	Binary	14 cols X 500 rows	Header	Hist	Plot	All	Sele
□ 7	AMBER_DATA	Binary	16 cols X 1500 rows	Header	Hist	Plot	Ali	Sele

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Tuesday June 16, 2006 18

# **OI-ARRAY** Table

lect All rert	_  TEL_NAME 8A	_  STA_NAME 8A	_ STA_INDEX I	_ DIAMETER E m	☐ STAXYZ 3D m Expand	
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2	UT3	<b>U</b> 3	33	8.000000E+00	Plot	
3	UT4	U4	34	8.000000E+00	Plot	

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# **OI-TARGET Table:**

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# OI\_WAVELENGTH TABLE

Select	⊔ EFF_WAVE E	⊔ EFF_BAND E	
	m	m	
Invert			
1	2.106147E-06	6.829834E-10	
2	2.106830E-06	6.829834E-10	
3	2.107513E-06	6.829834E-10	
4	2.108196E-06	6.828613E-10	
5	2.108879E-06	6.828613E-10	
6	2.109562E-06	6.829834E-10	
7	2.110245E-06	6.829834E-10	
8	2.110927E-06	6.828613E-10	
9	2.111610E-06	6.828613E-10	
10	2.112293E-06	6.828613E-10	
11	2.112976E-06	6.828613E-10	
12	2.113659E-06	6.828613E-10	
13	2.114342E-06	6.828613E-10	

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