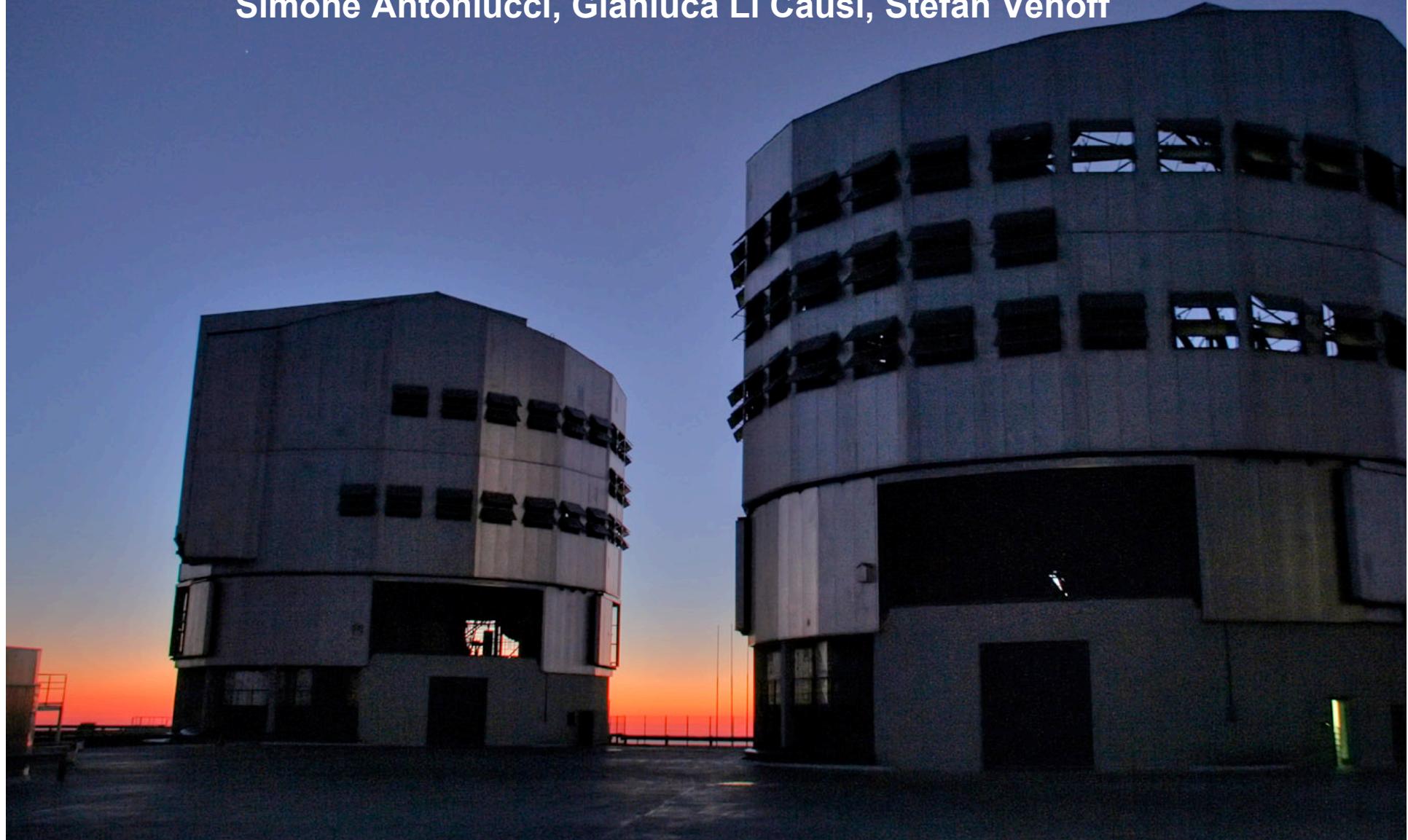
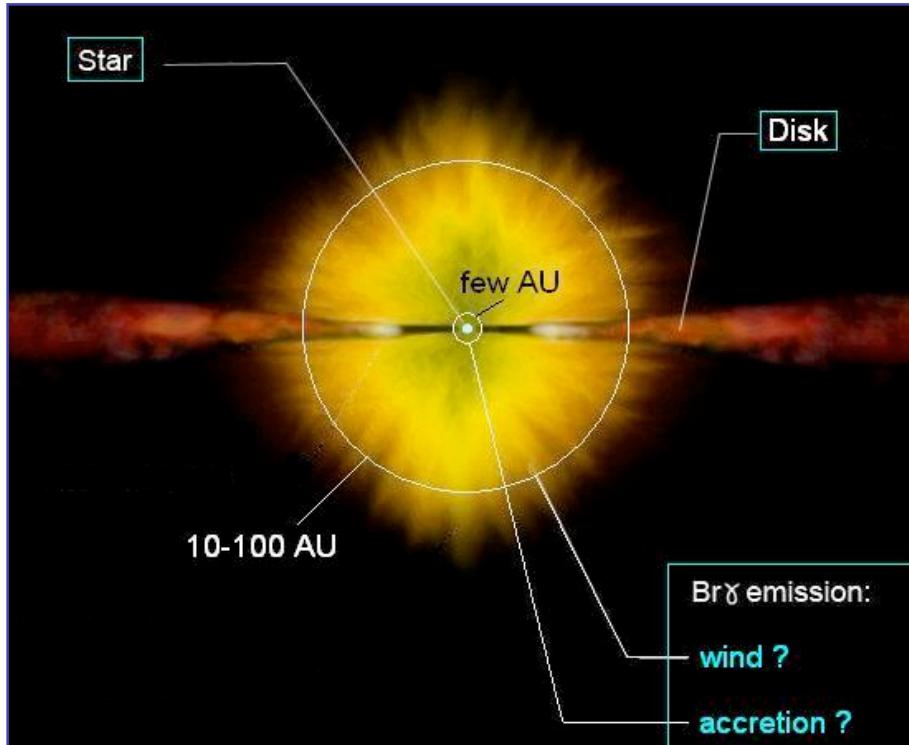


Studying the HI emission region around YSO NGC2024-IRS2

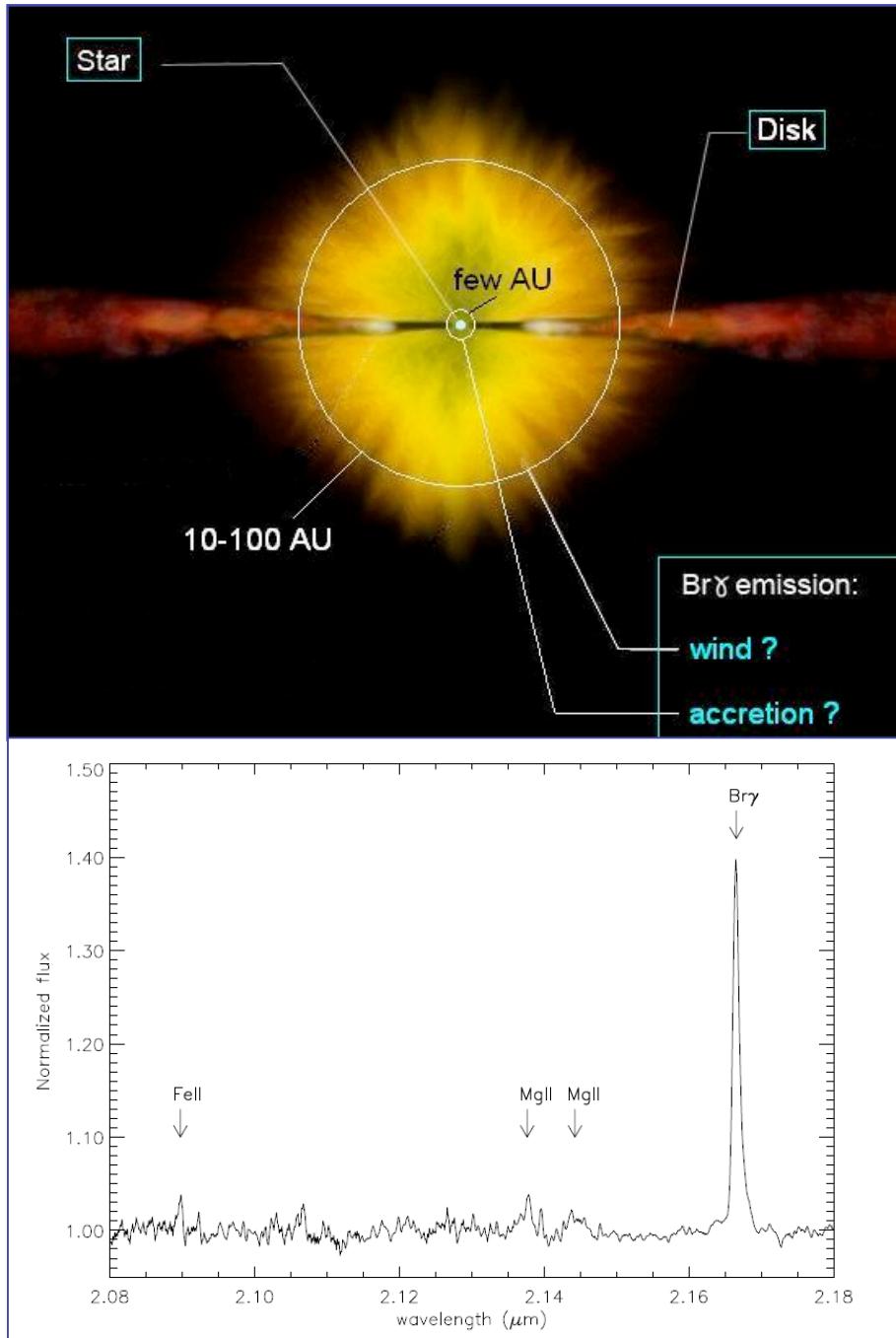
Simone Antonucci, Gianluca Li Causi, Stefan Vehoff





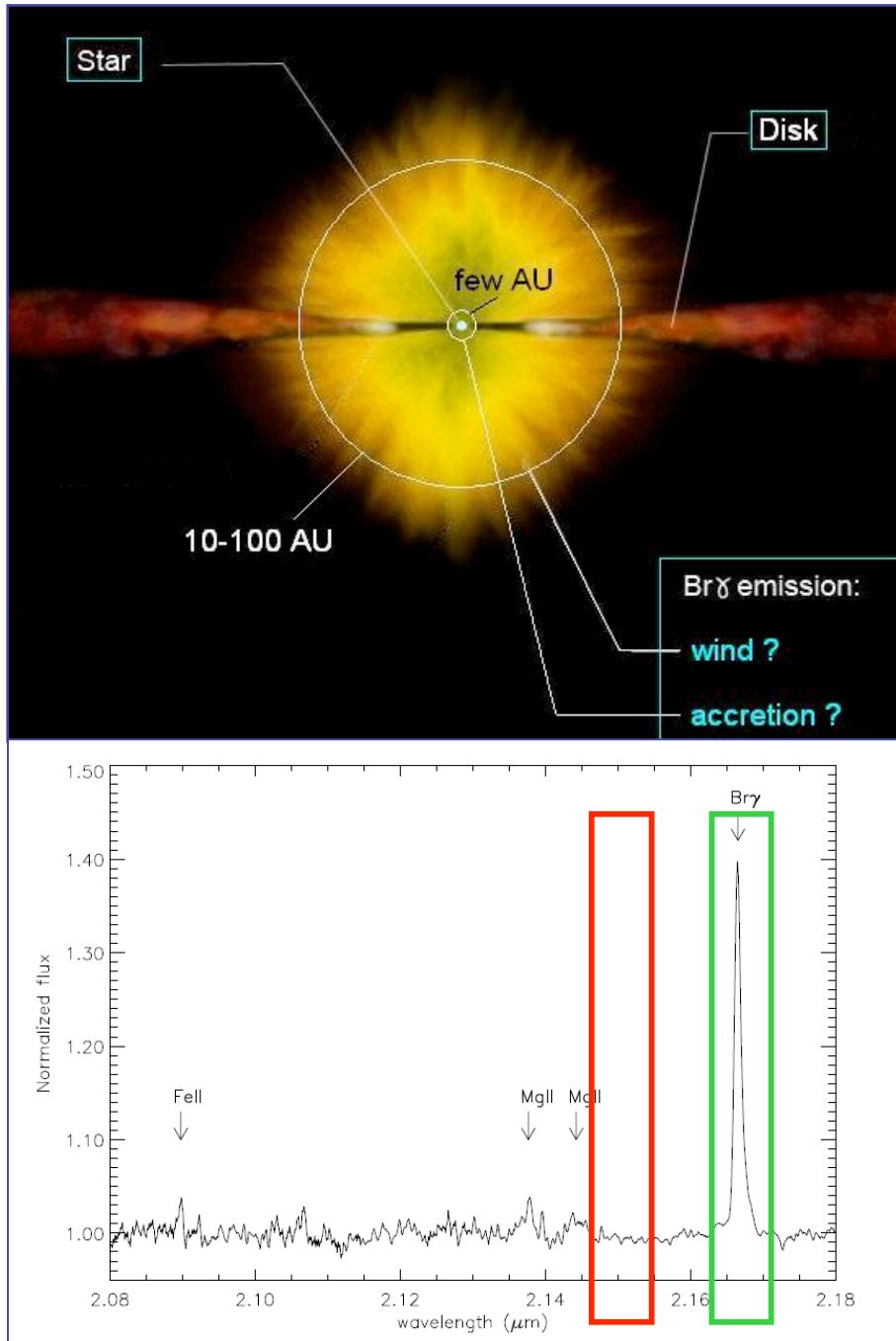
The object

- Class I YSO, $m_K = 4.5$,
 $d = 360$ pc, $L > 10^4 L_\odot$
- young star surrounded by an accretion disk; strong HI emission lines (e.g. Br γ)
- fit of SED yields disk with an outer radius of only 1AU
(Lenorzer et al. 2003)



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Investigating the inner region of the object with **AMBER**:

- check Lenorzer's results
- put constraints on the size of the emitting region: **wind** or **accretion flow**?

Preparing observations...

Target: NGC 2024-IRS2

Magnitude: 4.5 (K band)

VLTI configuration: UT1, UT2, UT4, min. 2 observations (4h)

Time of observation: late December, visible all night long

Configuration of AMBER: K Band – MR, spectral window centered around $2.17\mu\text{m}$

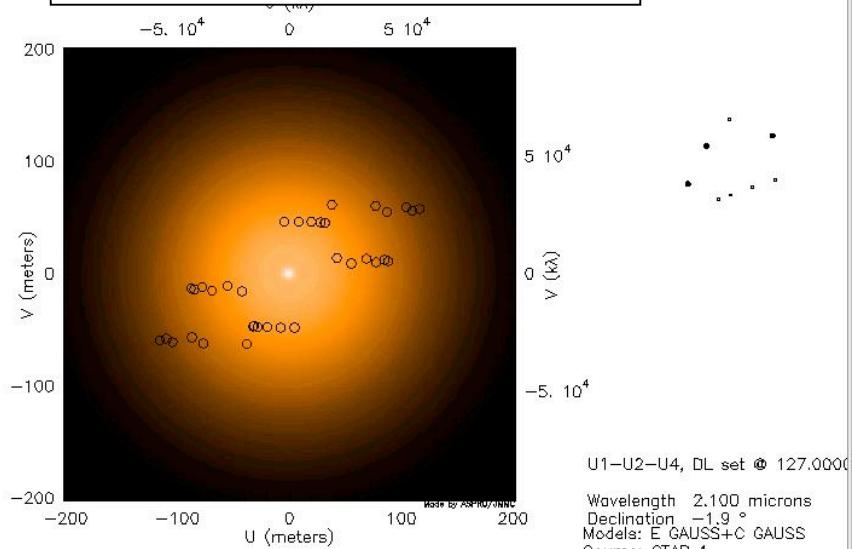
Calibrator: HD 30836

Magnitude: 4.1 (K band)

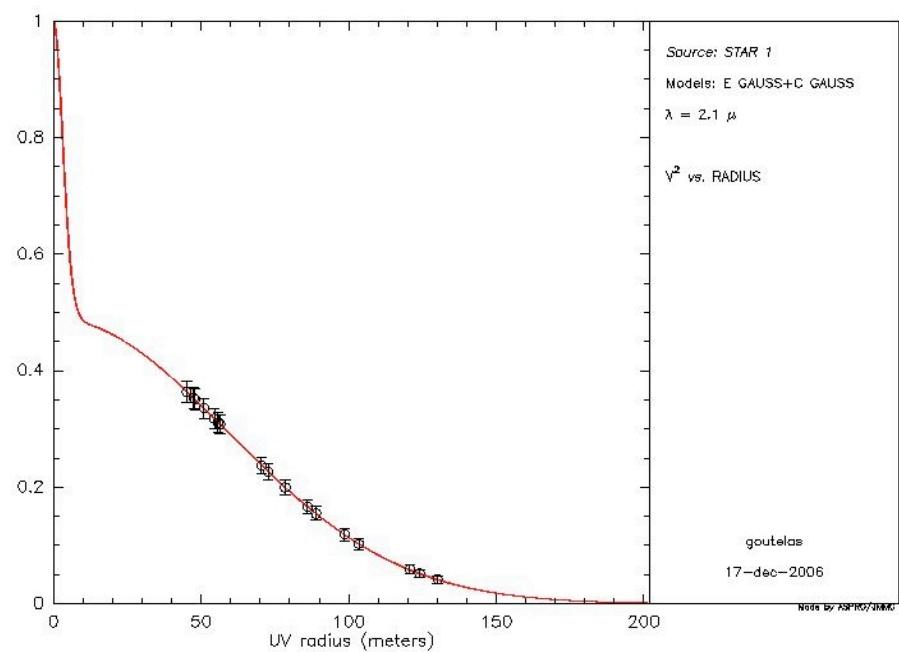
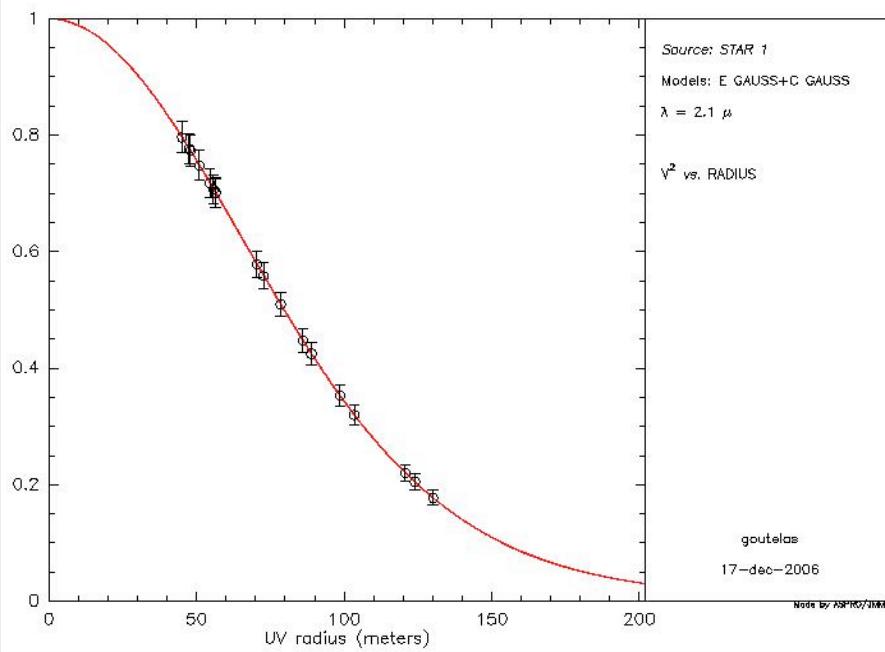
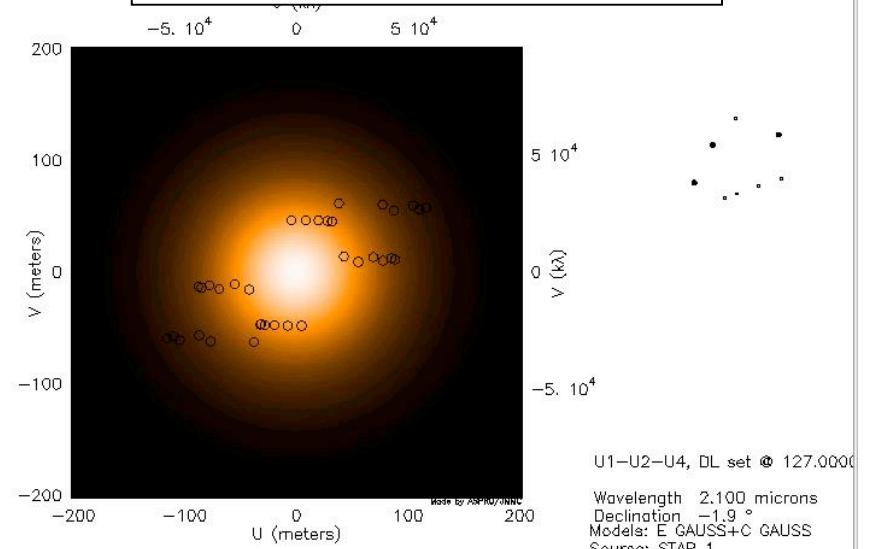
Diameter: 0.3 mas

ASPRO	Continuum emission (disk)	Bry emission (envelope)
ACCRETION Bry emission from small inner region	elliptical gaussian <i>major axis: 1.96 mas</i> <i>minor axis: depending on</i> <i>inclination</i>	circular gaussian <i>1 mas</i> <i>0.4 flux ratio</i>
WIND Bry emission from extended envelope		circular gaussian <i>50 mas</i> <i>0.4 flux ratio</i>

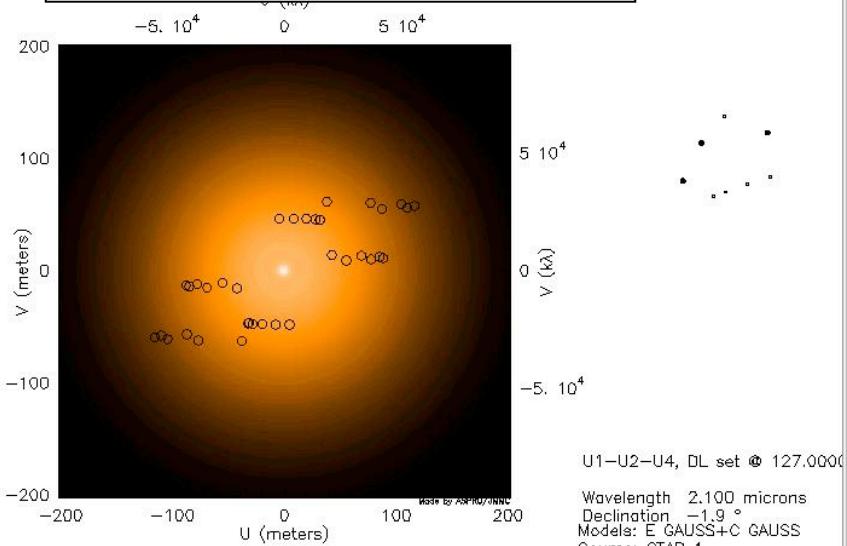
Bry, Accretion, pole-on



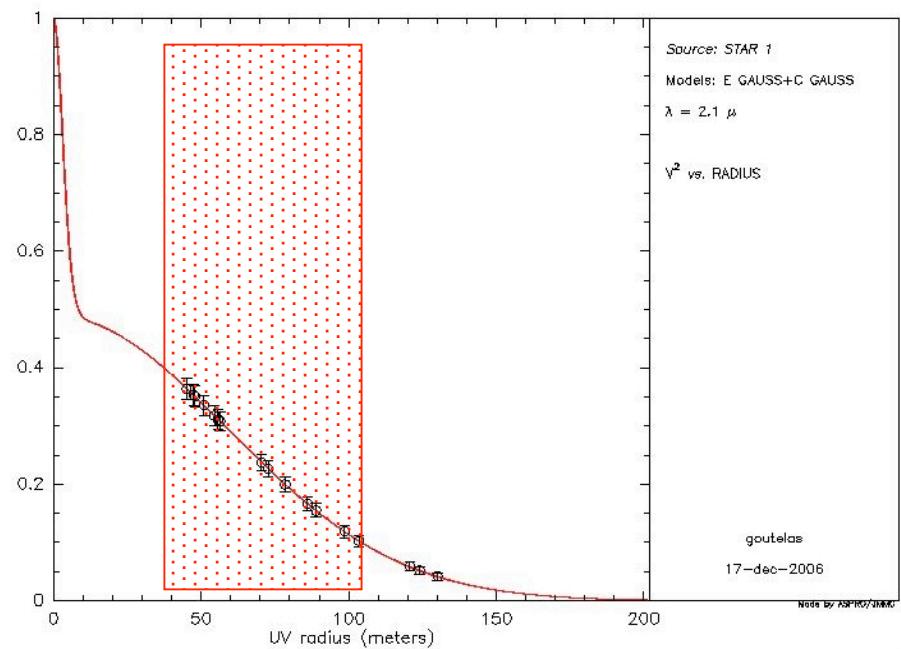
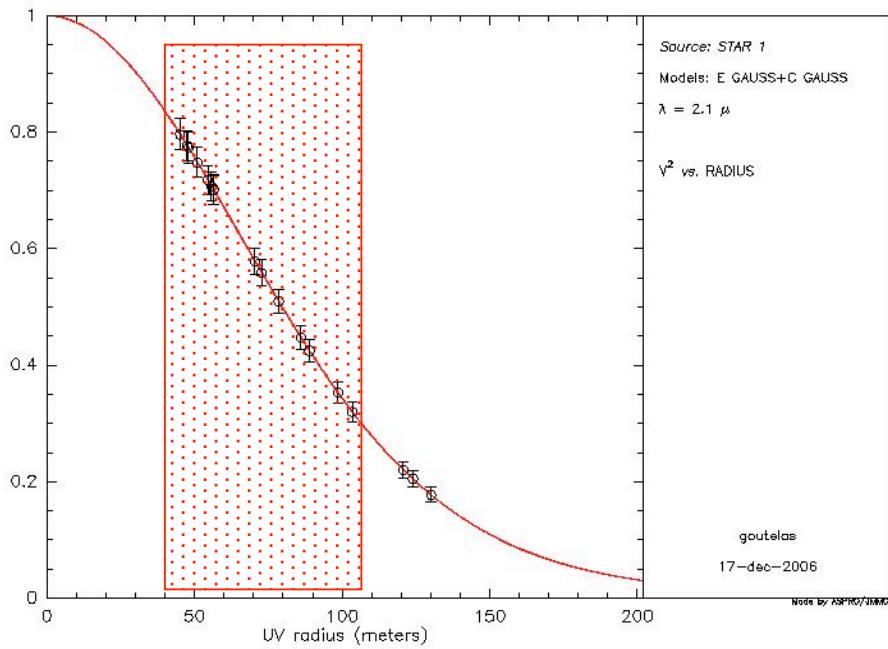
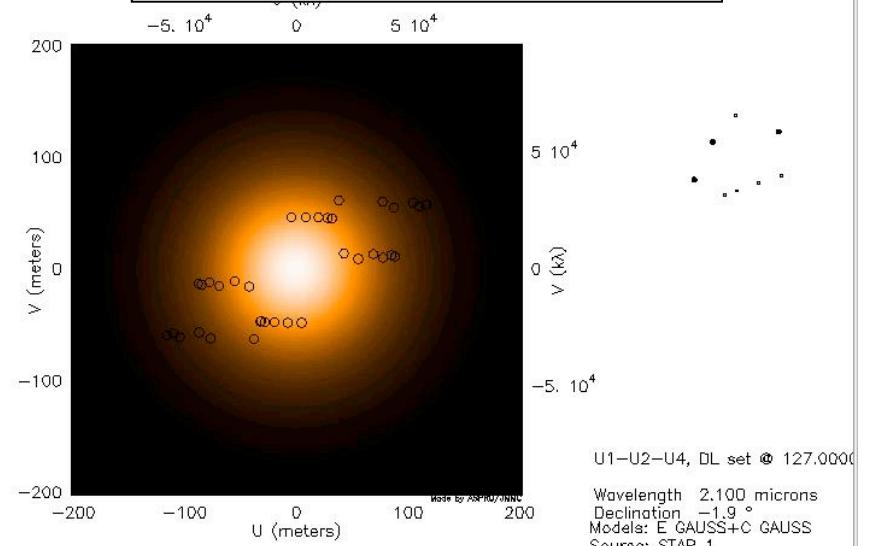
Bry, Wind, pole-on



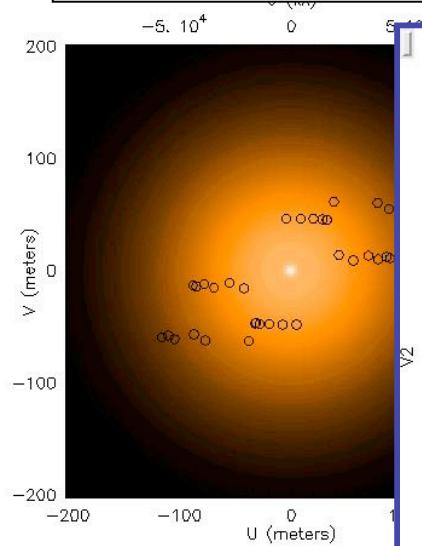
Bry, Accretion, pole-on



Bry, Wind, pole-on

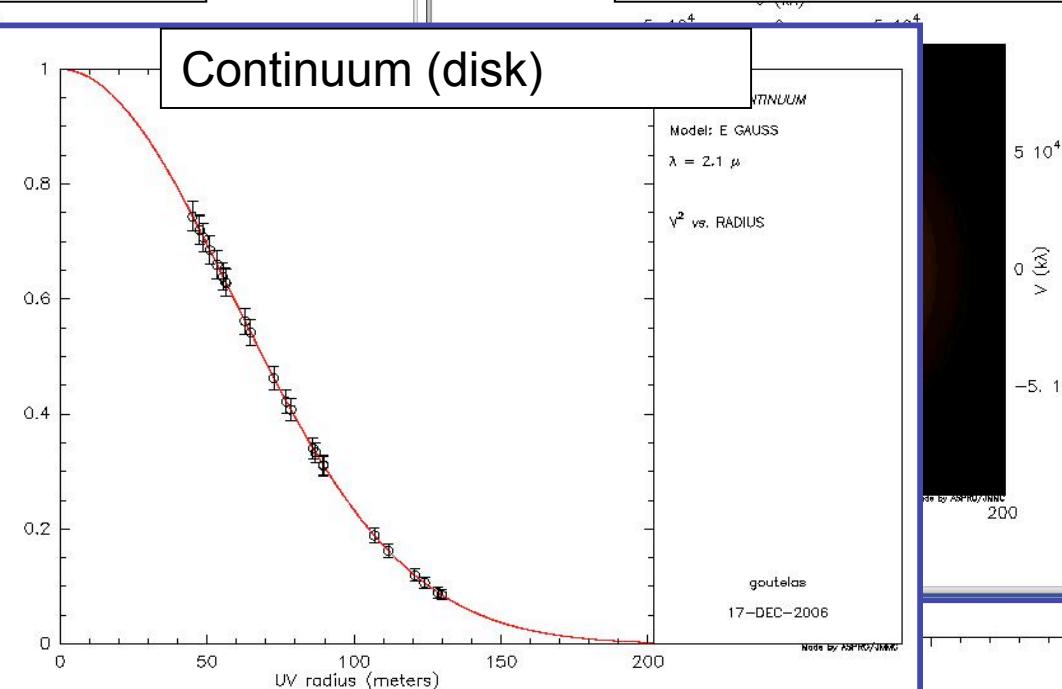


Bry, Accretion, pole-on



Bry, Wind, pole-on

Continuum (disk)



U1-U2-U4, DL set @ 127.0000
Wavelength 2.100 microns
Declination -1.9°
Models: E GAUSS+C GAUSS
Source: STAR 1

