EVOLUTION?





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poorly known

Bipolar morphology
Binary system separation not much larger than 10
AUs needed (models)

Case: OH 231.8+4.2, suspected but not resolved binary system

 Precession of molecular outflows → presence of companion
 Binary system? TiO & VO band → presence of a cool object M9-10 Mira Balmer series absorption → presence of a hot object A0



C. Sánchez Contreras, A. Gil de Paz, R.Sahai, 2004, ApJ, 616, 519-524

Observational setup



Calibrators

	Close Help													
		ſ	RESET SHOW ALL RES			ULTSSS	HOW DETAILS	HIDE	DETAILS					
01	Science star													
NAME	RAJ2000	RAJ2000 DEJ2000		MagK Base		Lambda								
OH231.8+4.	H231.8+4.2 07:42:16.83		9.470 102		45	2.16							-	
	Results													
Number of st	Number of stars: 31 found, 20 with coherent diameter and 4 without variability and multiplicity													
Number	dist	HD RAJ	<u>,2000</u> Г	DEJ2000	vis2	vis2Err	diam_vk	e_diam_vk	SpType	٧	T j	Н		
1	0.900 67	052 07 4	5 561	.4 33 4	0.058	0.006	0.574	0.040	F2V	5.036	4.570	4.328	4.2	
2	2.069 60	1552 1 07 3/	4 281	.3 52 1 🧹	0.984	0.002	0.349	0.024	F711/111	6.697	5.791	5.560	5.5	
3	2.438	1215 073	4 131	6 11 1	0.970	0.005	0.416	0.029	G2V	6.725	5.601	5.331	5.2	
4	4.850 65	846 08 00	0 101	.6 57 1	0.962	0.005	0.546	0.038	KOIII	7.022	5.385	4.990	4.7	

We have chosen the calibrator with the better compromise between the visibility and its error

uv plane coverage



Three points will constrain the model, but we have too many "free" parameters, so we apply for 6 visibilities per baseline.

Visibilities



Under our assumptions on the object we could resolve the binary system in one night 6h integration time.
The errors in the model fitting wil be: ~2% in both parameters: angular separation and Position Angle.