ASPRO2: A MODERN TOOL TO PREPARE OPTICAL INTERFEROMETRY OBSERVATIONS

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The Jean-Marie Mariotti Center is a network of French experts in optical interferometry and software engineers that develope, produce, document and maintain the software necessary for the exploitation and the followup of interferometric equipments, in particular the ESO/VLTI instruments.

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ASPRO2

ASPRO2 is the second version of the Astronomical Software to PRepare Observations created by the JMMC. It is quickly replacing its predecessor, ASPRO.

NON NOVA SED NOVAE: A NEW, SIMPLER, CLEARER USER INTERFACE



Feature	Aspro	Aspro 2	Comments
Dynamic Single Frame Interface	No	Yes	Plots updated on the fly without blocking user inputs
Standalone	No	Yes	Aspro2 does not depend on the JMMC server
Use Remote file transfer protocol	Yes	No	Aspro requires an account on the JMMC server to exchange files
Load / Save Settings	No	Yes	Useful to save your work and use it later off line or send it to collaborators
Export plots	Yes	Yes	PDF format
		Obje	ct definition
List of objects	No	Yes	
CDS Query	Yes	Yes	Both use SimBad service
Manual edition	Yes	Yes	Only coordinates in Aspro 2
Analytic models	Yes	Yes	Aspro 2 Model editor uses a GUI similar to LITpro and associates a model per object
User-provided model	Yes	No	Fits or 3D Data cube (multi wavelengths)
		Oh	servahility

ASPRO was a complete observation preparation tool that allowed to prepare interferometric observations with the ESO/VLTI. ASPRO was based on a client-server model, with a light java display interface on the client side and a complex server side, relying on a special "network-aware" version of the GILDAS software suite, a series of FORTRAN and C programs and SIC scripts.

Initially intended as a demonstrator only, ASPRO had a long and useful life (10 years), but is now difficult to maintain and improve due to its dependency to obsolete components. Based on the numerous positive returns and evolution requests from the community, the JMMC Scientific Council started in september 2009 the revewal of ASPRO...

Enters ASPR02

ASPRO 2 is a Java standalone program improving on all the functionalities of ASPRO and adding a dynamic graphical interface, the ability to use it off line, load and save observation settings, generate Observing Blocks, and much more. ASPRO2 is developped in close relationship with a panel of users, in a AGILE-like development environment.

ASPRO2 – GUI INTERFACE



Observabi bservability of Yes Yes oservability limits Yes Yes due to delay lines Yes Yes **Observability limits** CHARA nd best PoPs No Yes UV Coverage Aspro 2 does not support the Yes Yes superposition of several baseline configuration **UV Plots** Includes noise modelling Yes Yes **OIFits format Advanced features** Yes No Includes AO, FT ... Knows about nagnitudes arch Calibrators Yes (No) Use SearchCa lodel Fitting Use LITpro (No) Yes IFits explorer No To Be Done Yes VLTI AMBER / MIDI No Yes Export to SO OBs Compatible with P2PP No CHARA VEGA Yes xport to

Compatible with VEGA_PLAN

UV COVERAGE Based on :
Object's observability
Object's analytical model
Instrumental configuration

ASPRO2 – THE COMPONENTS

 JMCS by S. Lafrasse, G.Mella (JMMC). Shared library providing GUI and common features.
 Aspro2 core library

 JskyCalc 1.2 by J. R. Thorstensen, Dartmouth College: ephemeris, astronomical coordinates conversions, etc: the only JAVA library available
 nom.tam.fits by Dr Thomas A. McGlynn, HEASARC : FITS file handling. We added the to support of single and double-precision Complex values and handling of the COMMENT and UNIT keywords

- JFreeChart vector plots. Exports in SVG or PDF.
- JSAMP 1.1 by Mark Taylor, AstroGrid : for the SAMP VO query protocol.

 SIMBAD Service @CDS to retrieve the relevant information (position, magnitude, proper motions, etc...) simply by entering the object's name.

ASPRO2 – The Data Model

ASPRO2 is based on a Data Model (xml configuration files) to easily maintain and update the configuration for interferometers, instruments and observations:

The interferometer itself Telescopes sizes, Stations positions,

OBSERVABILITY Based on:

- Interferometer configuration
- night restriction for the observation date
- chosen minimum elevation
- delay line compensation for the selected base lines
- telescope shadowing (VLTI), zenihtal constraints
- CHARA'S Pipes Of Pan (PoPs)



Geometrical delays & constraints





🔻 🚞 Models	Model actions
 HIP 100345 disk1 punct2 	Name Add Remove model type punct Update

Returns the Fourier transform of a punctual object (Dirac function) at coordinates (X,Y (milliarcsecond). FLUX_WEIGHT is the intensity coefficient. FLUX_WEIGHT=1 means total energy is 1.

Model	Name	Units	Value
disk1	flux weight1		0.9
	×l	mas	0
	yl	mas	C
	diameterl	mas	3.45
punct2	flux weight2		0.1
	rho2	mas	42.426
	theta2	deq	45

OK Cancel

Optical Paths, Delay lines, Atmospheric conditions, etc... example: CHARA, VLTI...

The instrument used Transmissivity, bandwidth, number of spectral channels and resolution, detector properties... example: AMBER, MIDI, VEGA, CLIMB, CLASSIC, MIRC.

The instrument noise At this time we have an all-purpose "generic" noise model valid for any fibered recombiner (monoaxial/multiaxial).
 The Object based on a collection of simple parametric models. Exemple: point source, elliptical disk, gaussian...

We use the OI-FITS format to store the simulated observations. OI-FITS is the first attempt at a norm for data exchange of Optical Interferometry data and is widely used in the OI community.

REFERENCES

ASPRO:

Duvert G., Bério P., Malbet F., 2002, SPIE, 4844, 295
 Mella G., Duvert G., 2004, SPIE, 5496, 582
 SearchCal:
 Bonneau, D., Clausse, J.-M., Delfosse, X., et al, 2006, Astron. Astroph. 456, 789.
 LITpro:
 Tallon-Bosc, I., Tallon, M., Thiébaut, E., Béchet, C., et al., 2008, SPIE, 7013.



SearchCal searches calibrators and sends a VOTable back...

1) Ins	trumenta	l Configura	ation	2) Science Object 3) SearchCal Parar					ameters	neters						
Ma	Magnitude Band : Κ 🗸 🗸			Name : Q HIP32768 RA 2000 [hh:mm:ss] : 06:49:56.168) Min. Magnitude (K) : -1.664									
Wavel							Max. Magnitude (K): 2.336									
Max	. Baselin	e [m]: 102	43383	DEC 2000 [+/-dd:mm:ss] : 50:36:52.415 Magnitude (K) : 0.336			Scenario : Bright Eai									
							RA Range [mn] : 240.0									
								DEC Range [deg] : 20.0								
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A. T																
Found	Calibrato	rs														
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1	10.217	<u>63744</u>	07 48 20	47 04 39	0.659	0.0090	1.657	0.114	1.525	1.593	1.606					
2	10.509	39523	05 49 49	56 09 59	0.545	0.018	2.02	0.139	1.86	1.942	1.958					
3	16.668	73155	08 34 43	49 56 39	0.507	0.051	2.058	0.142	1.886	1.973	1.99					
4	21.038	14/12	08 44 23	42 38 57	0.607	0.044	1.778	0.123	1.652	1./1/	1.73					
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Aspro 2 displays returned calibrators in blue



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